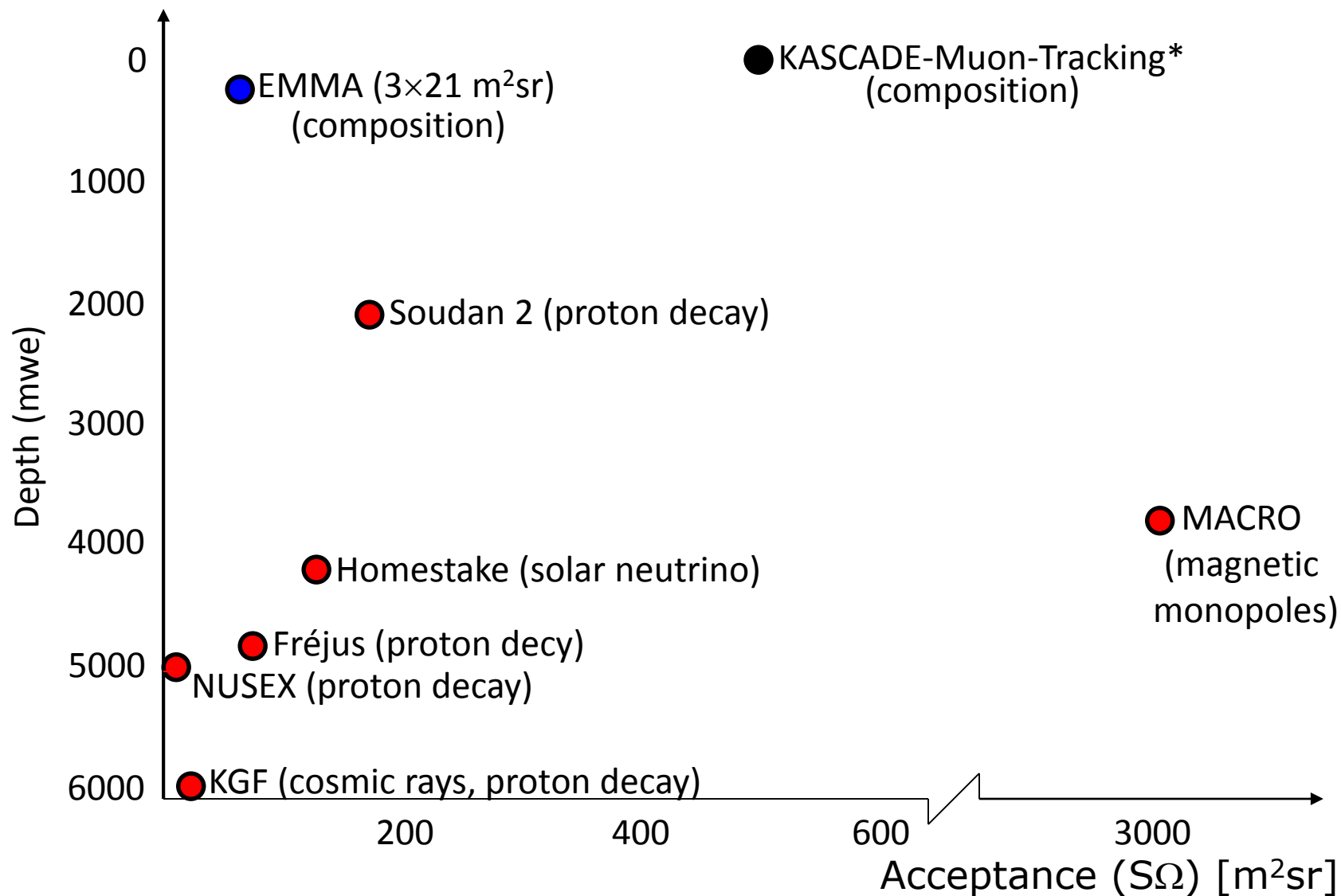




Muon multiplicities measured using an underground cosmic-ray array

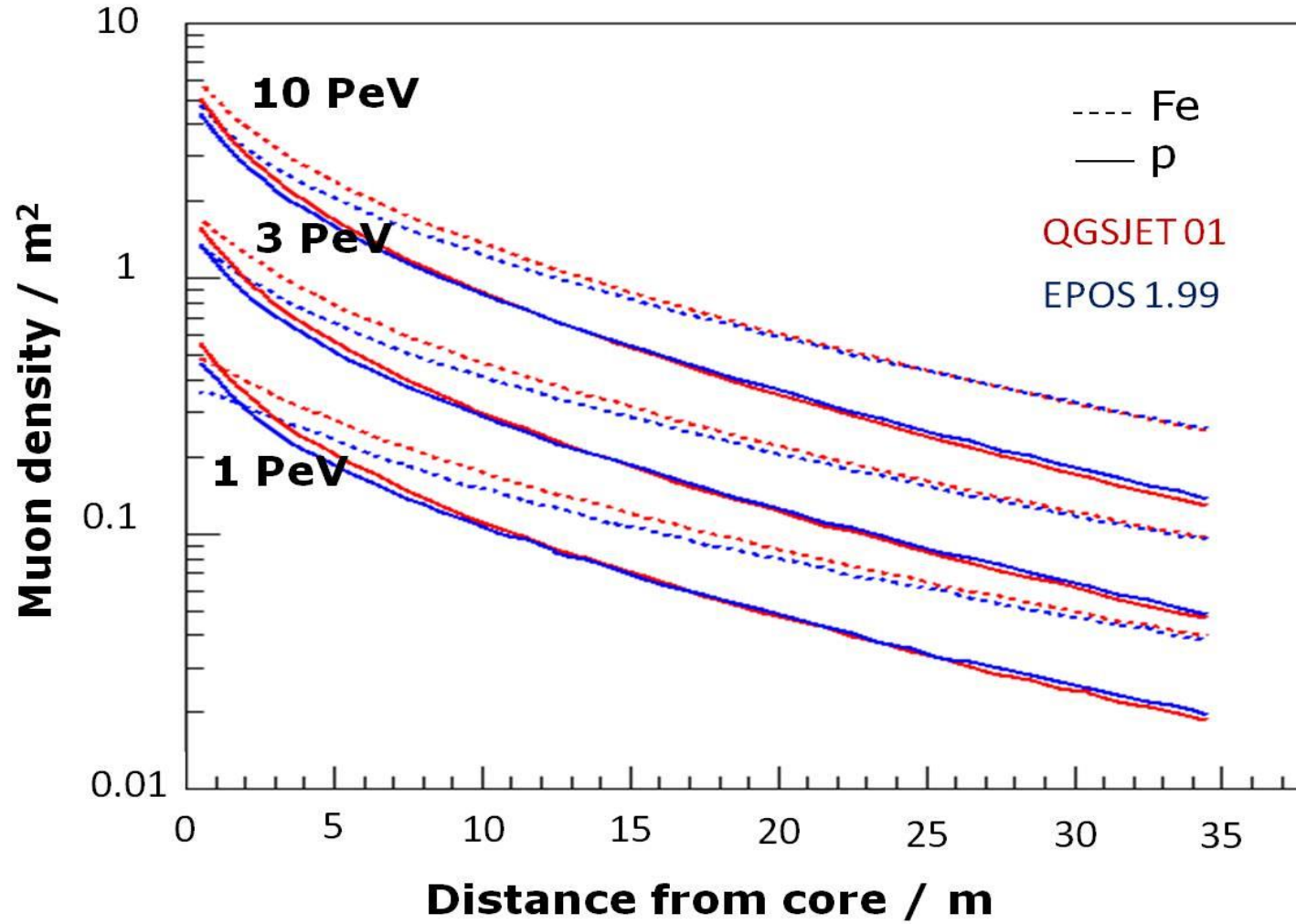
Pasi Kuusiniemi, University of Oulu, Finland
for the EMMA collaboration

- intro
- methods/examples
- comments/questions/discussion



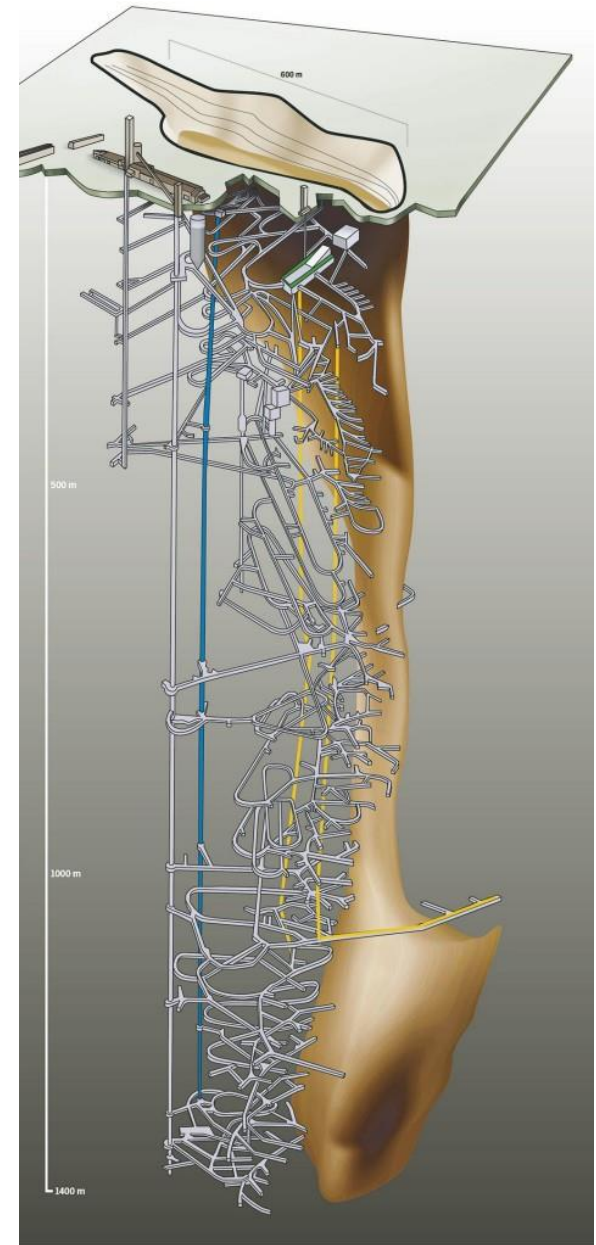
*The Muon-Tracking-Detector only, the full array is on ground and very large.

CORSIKA+QGSJET 01 and EPOS 1.99 – 50 GeV muon cutoff



- + muon density at the shower core \propto primary energy
- + gradient \propto mass

Pyhäsalmi mine (1.4 km)



+ opened 1962,

+ ore: copper, zinc and pyrite (gold, silver),

Experiment with Multi-Muon Array

- + designed for cosmic-ray composition studies around the knee energy (1 – 10 PeV),
- + measures air-shower initiated muon lateral distributions event-by-event \Leftarrow measuring the multiplicity and the lateral spread of high-energy muons in each shower,
- + operates 75 m below ground (210 mwe) in the Pyhäsalmi mine, Finland,
- + rock overburden \Rightarrow muon cutoff energy of ≈ 50 GeV (for vertical muons) \Rightarrow detected muons are mostly generated in the upper part of the air shower, or close to the primary interaction,
- + consists of 11 detector stations,
- + employs three different kinds of detectors,
- + muon tracking is used to extract the shower properties.

+ 11 stations

- 9 stations at 75 m-level and two at 45 m-level,
- area $\approx 15 \text{ m}^2$ each,
- tracking stations and sampling stations.

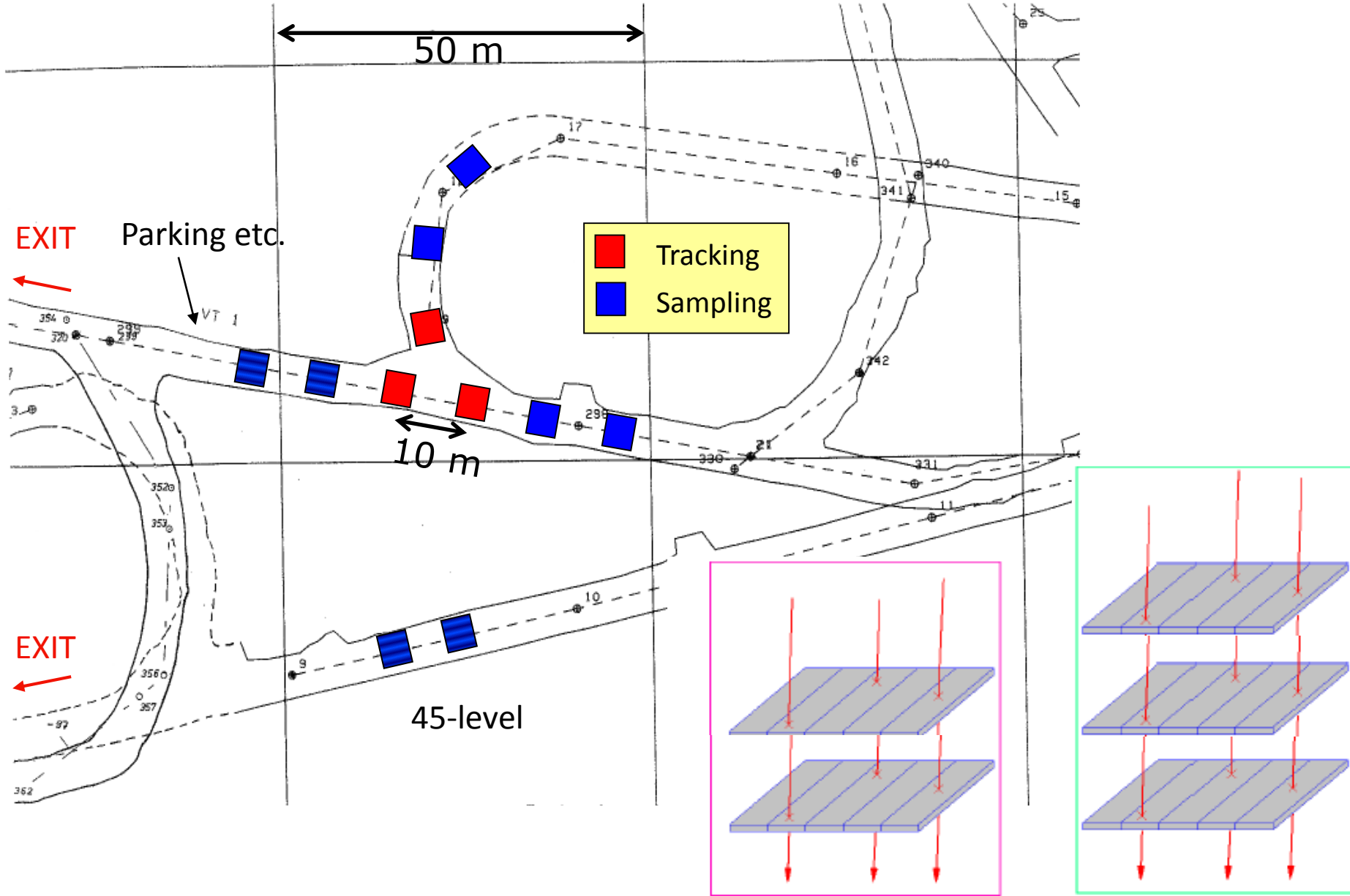
+ tracking stations

- use all three detectors types,
- muon multiplicities at the shower core (i.e. energy),
- arrival angles (position is very accurate, $\sigma \sim 1 \text{ cm}^2$),
- on-line efficiency calibrations (12h bins),
- tracking.

+ sampling stations

- find muon multiplicities (or densities) in the tails (i.e. mass).

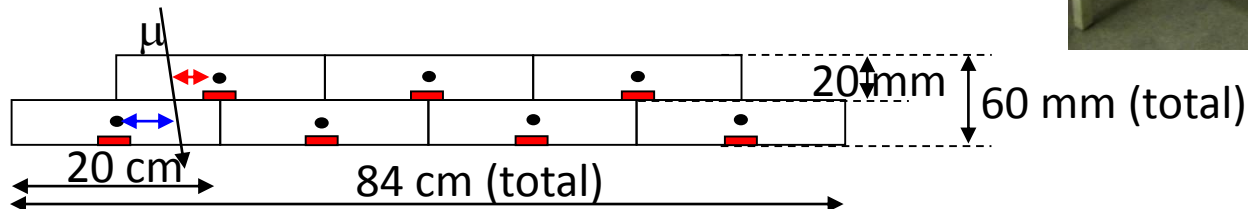
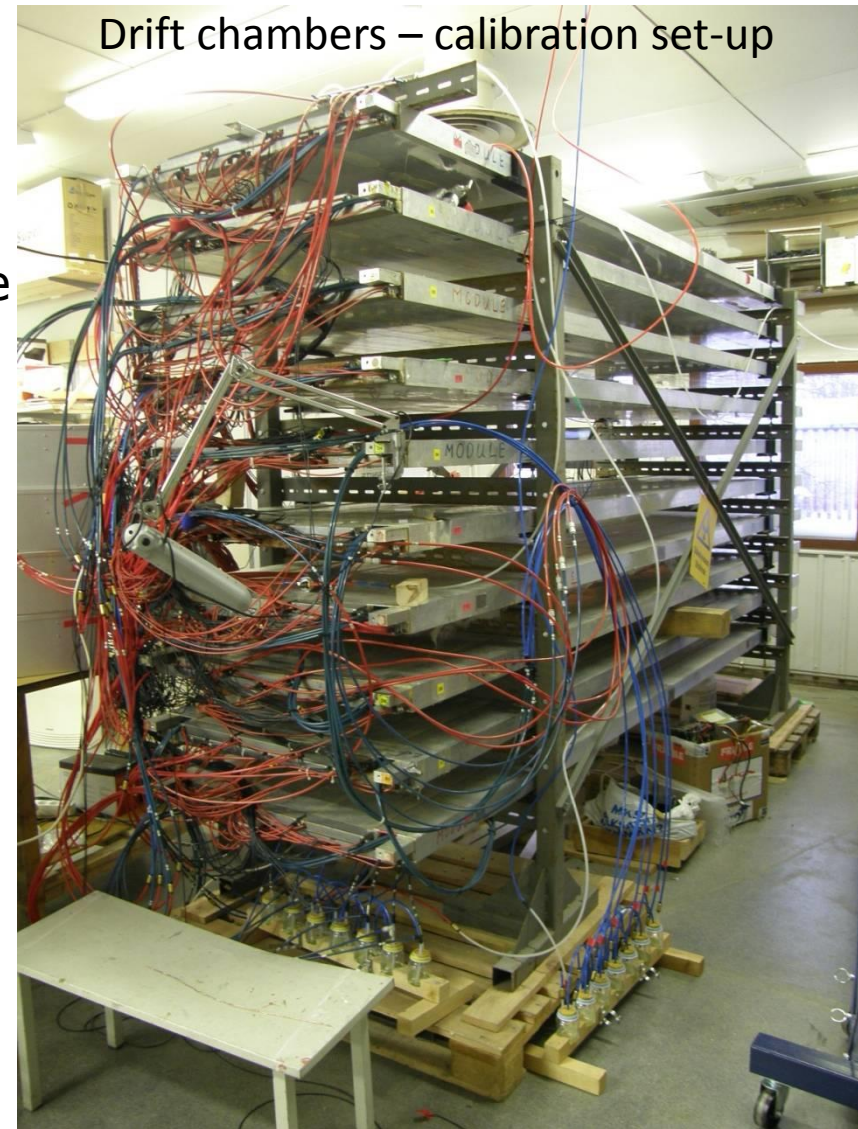
EMMA at 75m below ground





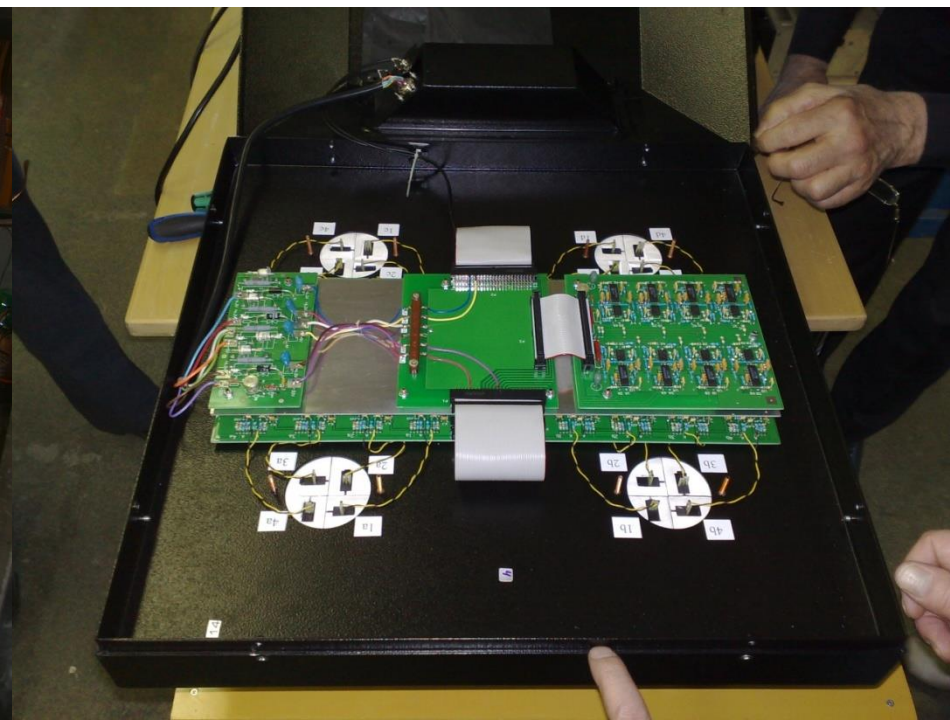
Former LEP-DELPHI MUBs (or planks)

- Ar(92%):CO₂(8%)-mixture delivered from ground via ≈ 100 metres long pipeline through the rock (gas consumption ≈ 10 l/min, NTP),
- seven position sensitive drift chambers (365×20 cm², 20 mm thick),
- arranged in lengthwise half-overlapping groups of 3 + 4 (area of 2.9 m²),
- position resolution (σ) $\approx \pm 1$ cm²,
- total area ≈ 240 m² (84 planks).



Plastic scintillation detectors (SC16s from INR/RAS, Moscow)

- designed for EMMA (muon multiplicity, timing),
- 4×4 individual 12×12 cm², 3 cm thick pixels (area of 0.5×0.5 m²),
- equipped with APDs (Avalanche Photo Diodes),
- total number of SC16s is 96 (or 1536 individual pixels, or 24 m²),
- placed in the three central tracking stations,
- time resolution (σ) better than 2 ns.



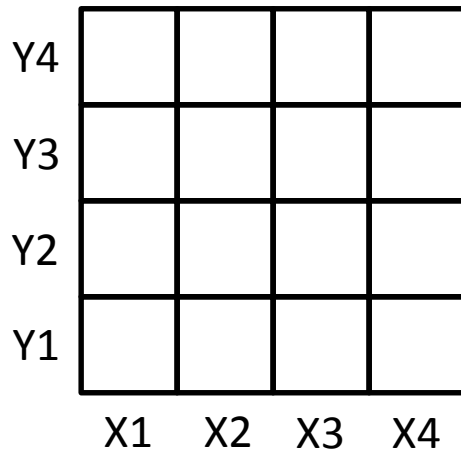


Station G – Mid level

Station F – Bot level



SC16-pixels



Limited Streamer Tubes (dismantled from KASCADE-Grande)

- gas-filled (with CO_2) position sensitive, area of 2.9 m^2 ($1.0 \times 2.9 \text{ m}^2$),
- the number of modules is 60 \Rightarrow total area $\approx 174 \text{ m}^2$,
- a pad (or pixel) size is $2 \times 8 \text{ cm}^2$.

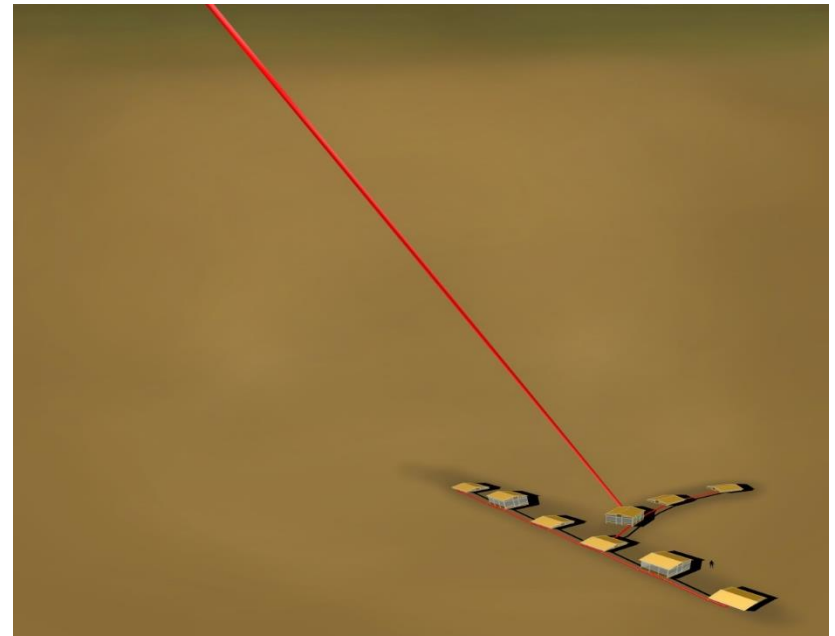
New read-out electronics designed and expected to be in operation early next year.





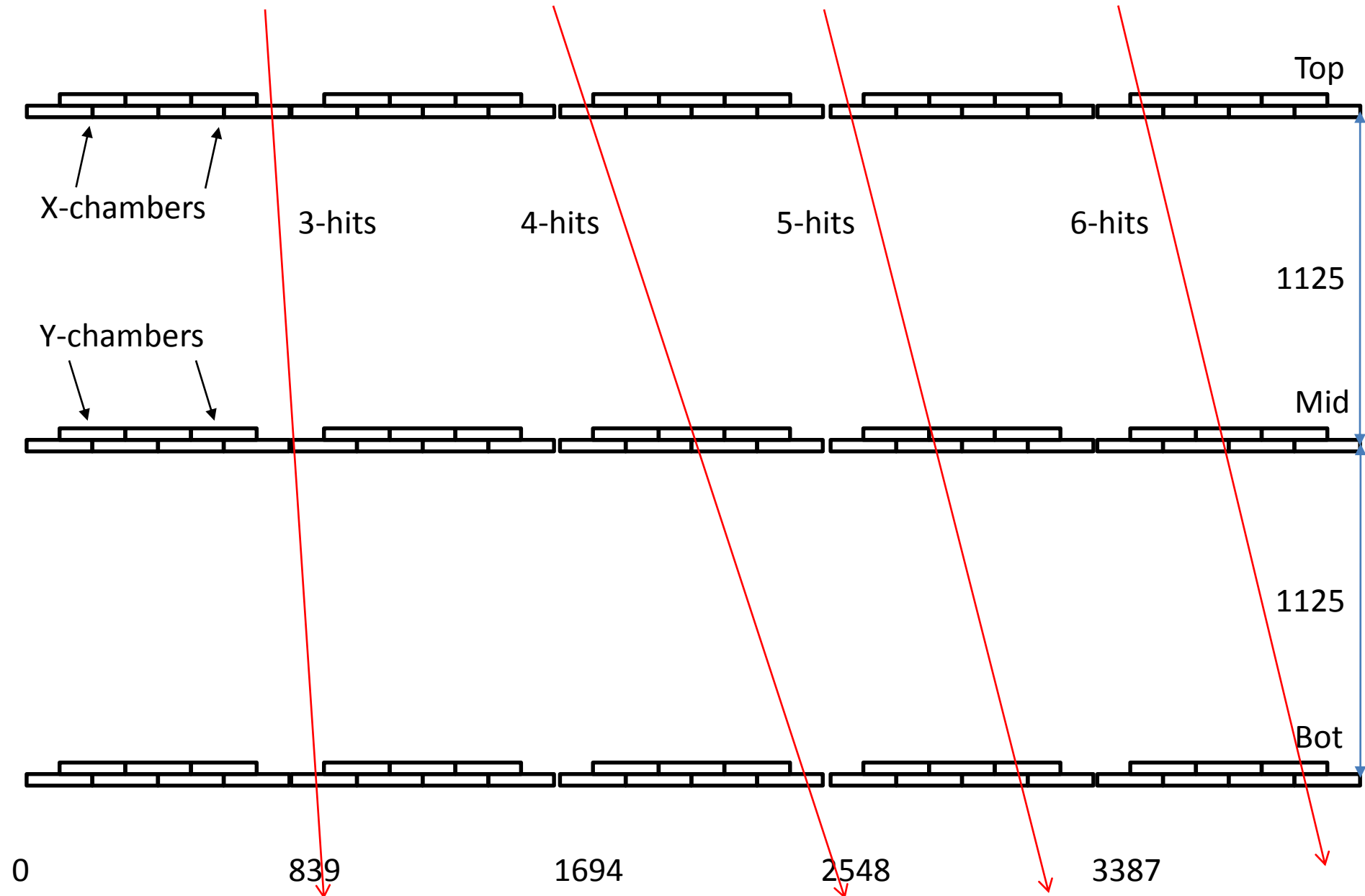
95 metre pipeline

- + Ar+CO₂-mixture for drift chambers,
- + CO₂ for LSTs,
- + Ar from LAr,
- + LAr and CO₂ in 3 m³ tanks,
- + gas consumption ≈ 10 l/min, NTP.



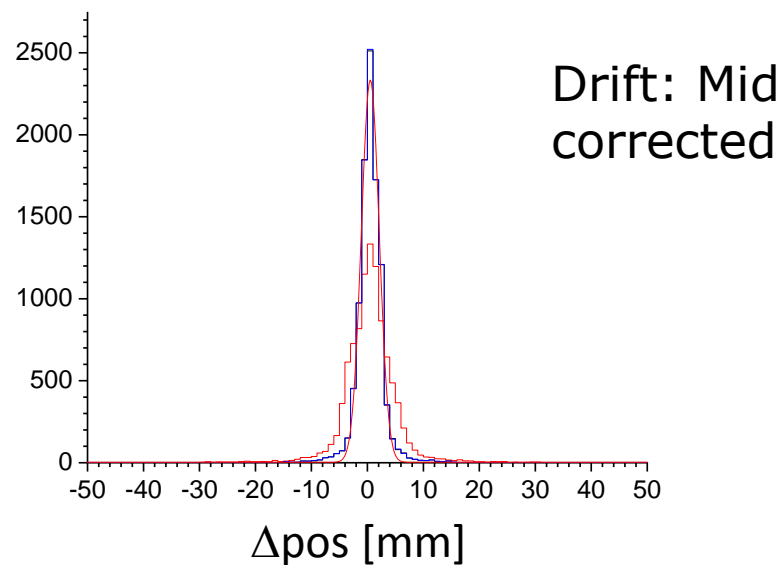
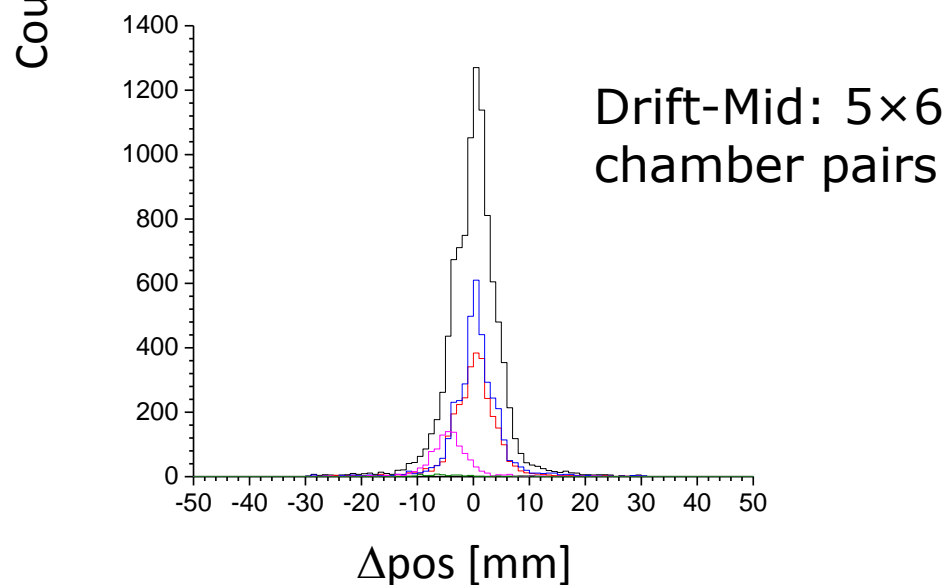
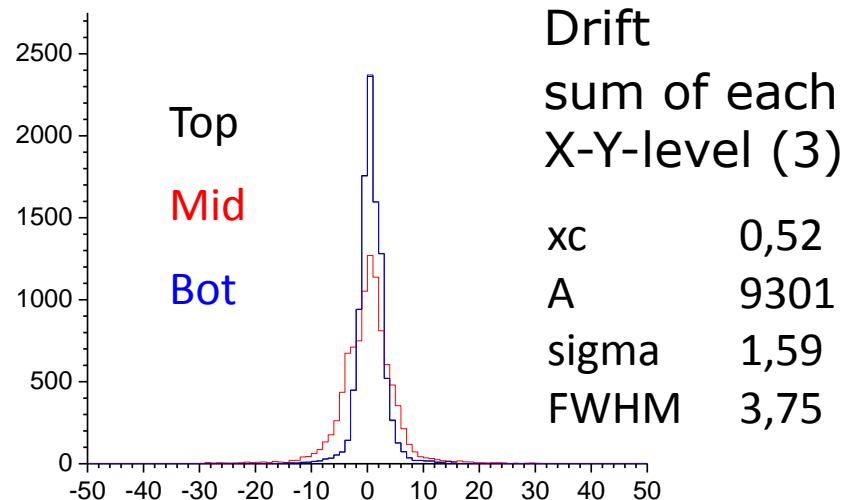
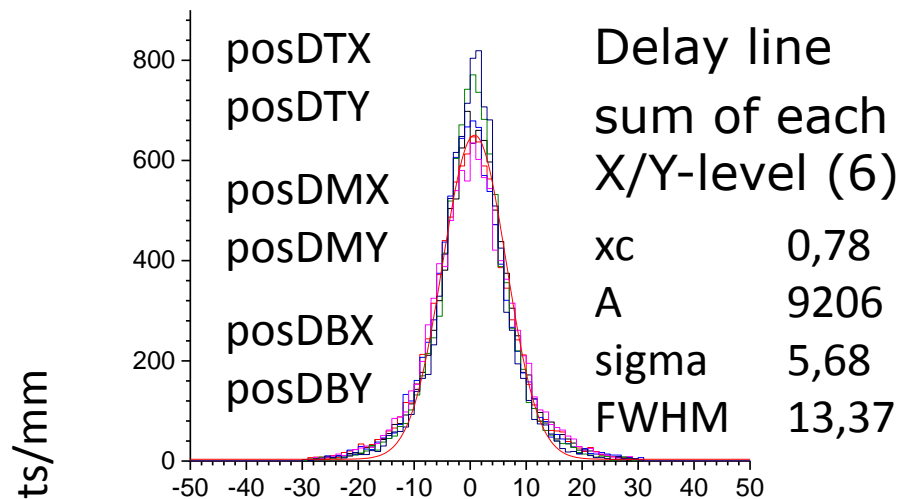


Tracking – possible tracks



Residual $\Delta\text{pos} = \text{pos}_{\text{track}} - \text{pos}_{\text{meas}}$

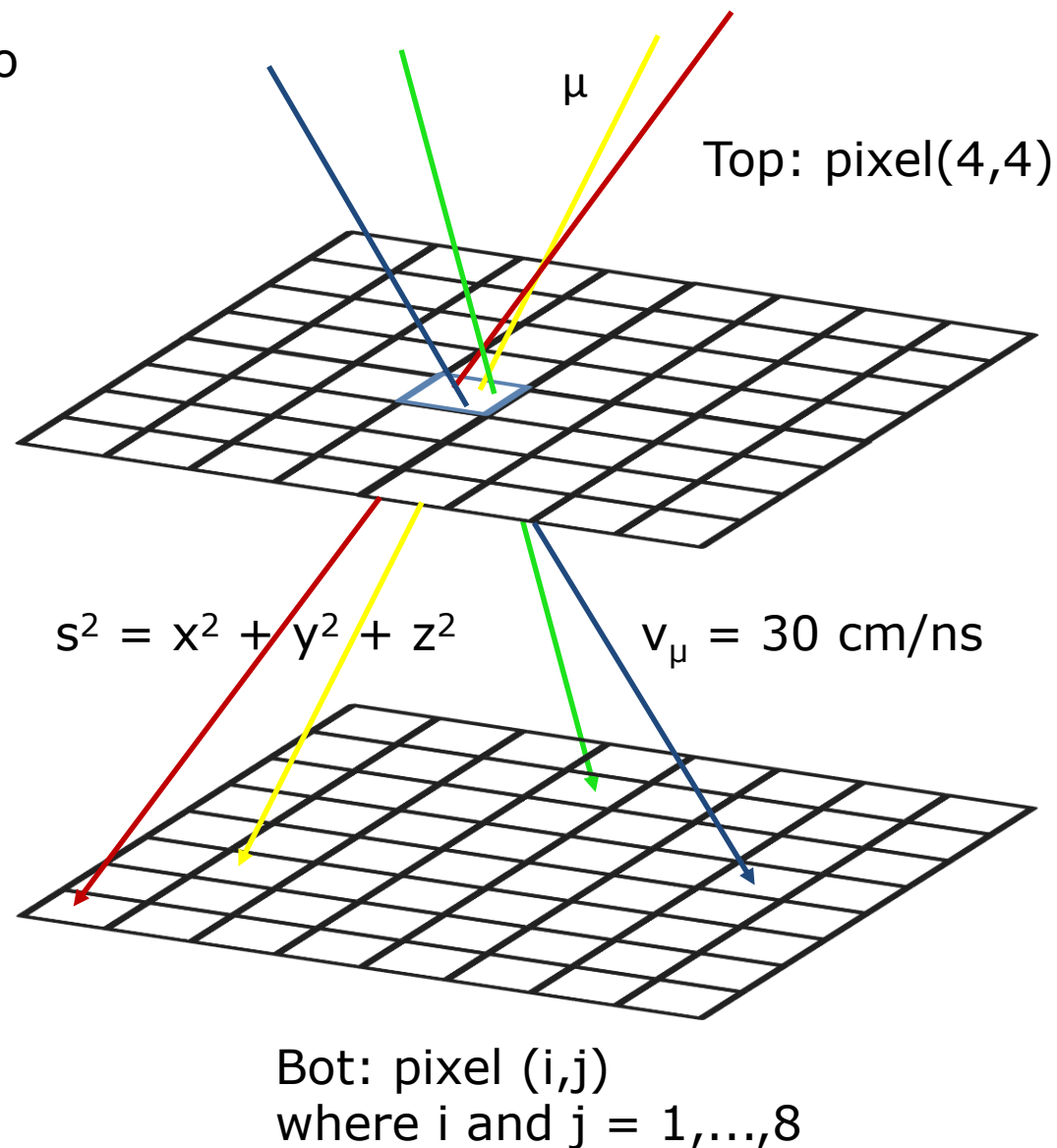
R246-F-12h



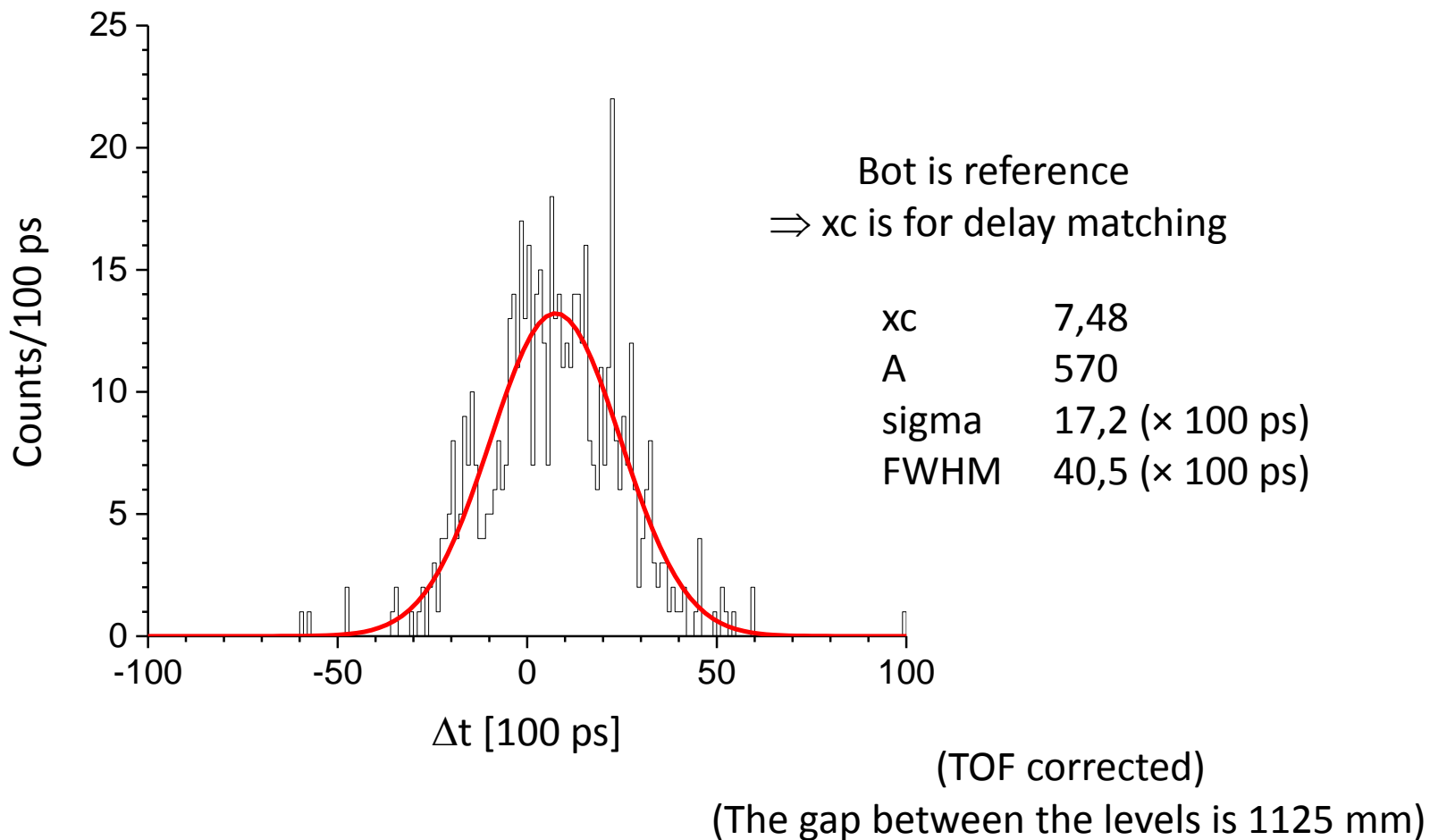
(use tracking and)

find coincs between (pixels in) two levels and correct for TOF.

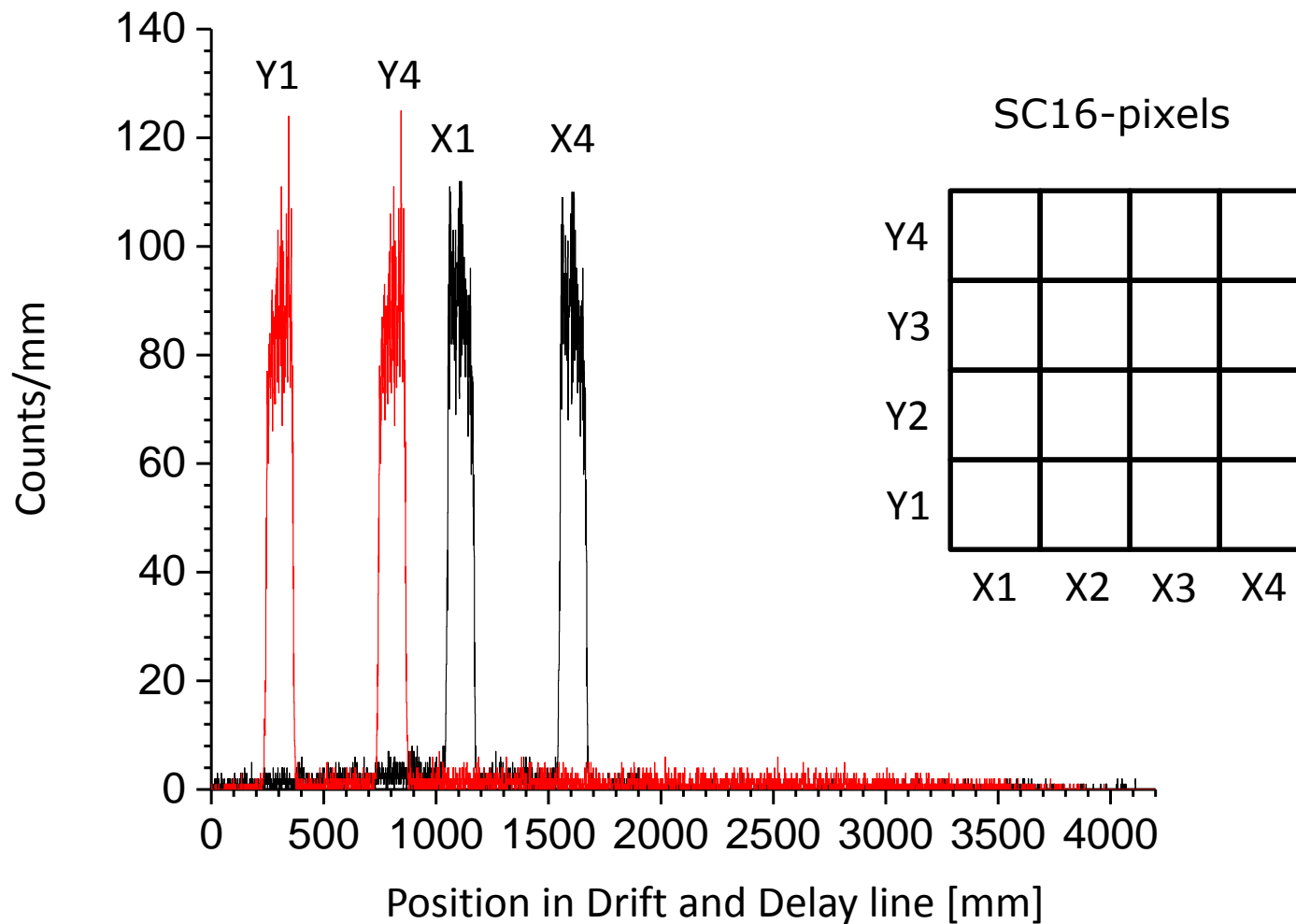
- muon multiplicities,
- fast timing,
- background reduction,
- initial guess for arrival angles,
- etc.



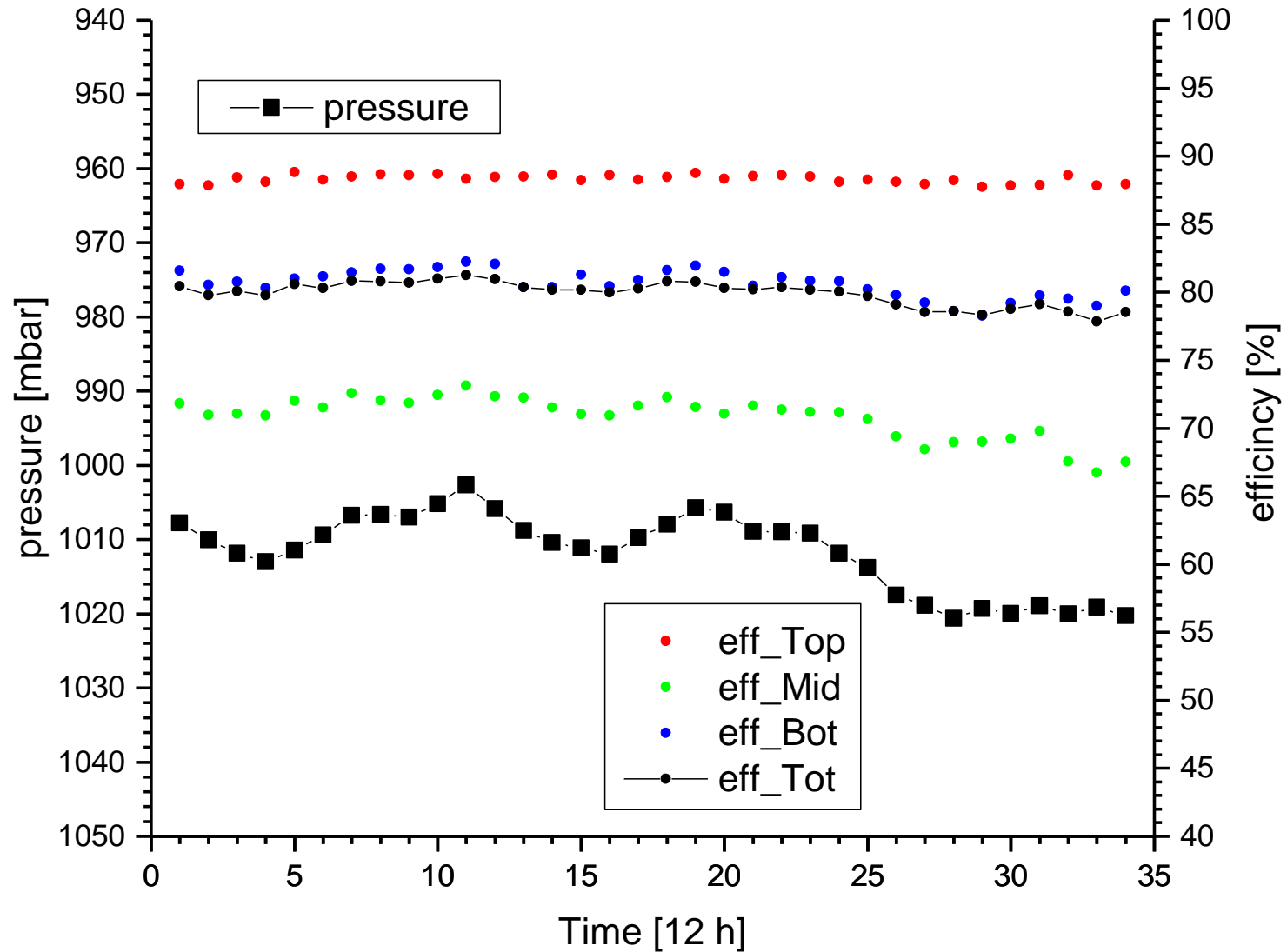
Time difference Δt ("one pixel on Top" – "8×16 on Bot") after delay matching, Station G, SC16 positions taken from tracking.



extracted using tracking, Station G



Station G, R246: 34x12h = 408h of data



Check efficiencies using single muon flux:

$$\text{Flux [Hz/m}^2\text{]} = (n_{\text{obs}} \times \text{eff}_{\text{detection}} \times \text{eff}_{\text{geometry}}) / (D \times \text{Area}_{\text{active}})$$

n_{obs} = the number of observed (single) muons,

$\text{eff}_{\text{detection}}$ = detection efficiency, e.g., for 6-hit track = $1/\sum e_i$ where e_i is the muon detection efficiency of chamber i ,

$\text{eff}_{\text{geometry}}$ = corrects for losses in detector area ("broken" chambers, gaps between the chambers),

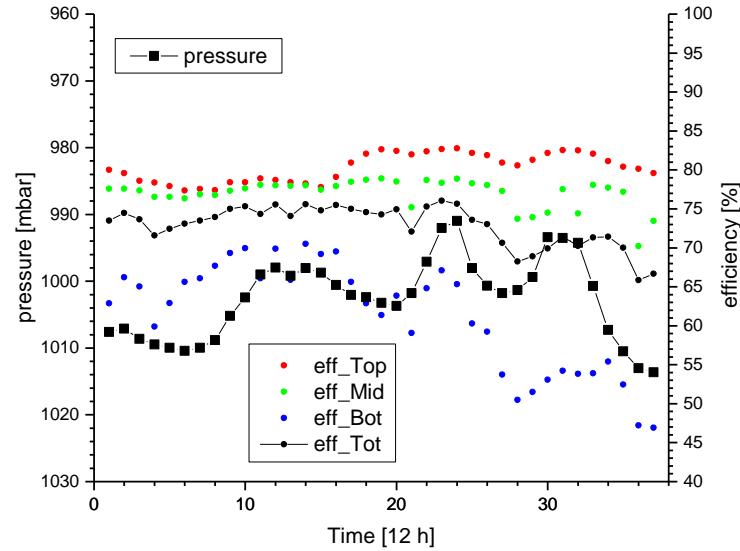
D = duration,

$\text{Area}_{\text{active}}$ = detector area.

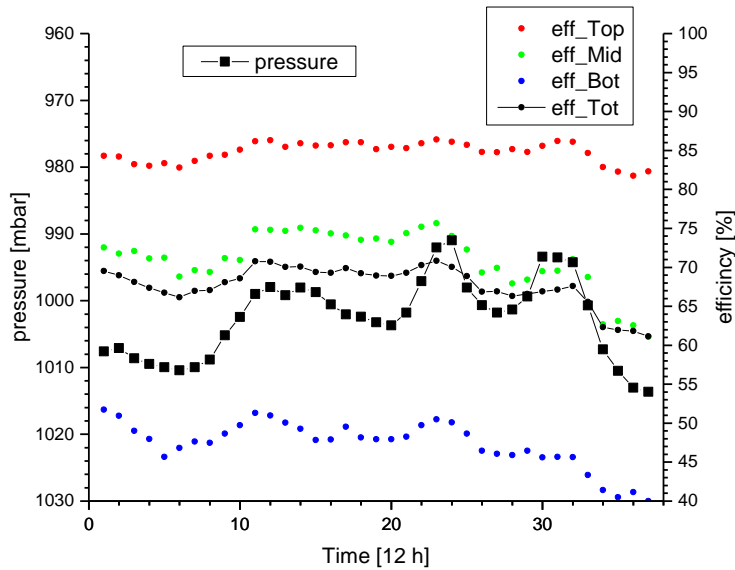
However, this is a simplified picture. In reality ' $\text{eff}_{\text{geometry}}$ ' requires assumptions on the muon distribution (through the rock) which is assumed here to be simple, or $\sin\theta\cos\theta$ where θ is the zenith angle.

⇒ here we simply use muon flux per steradian, or $[\text{Hz/m}^2\cdot\text{sr}]$.

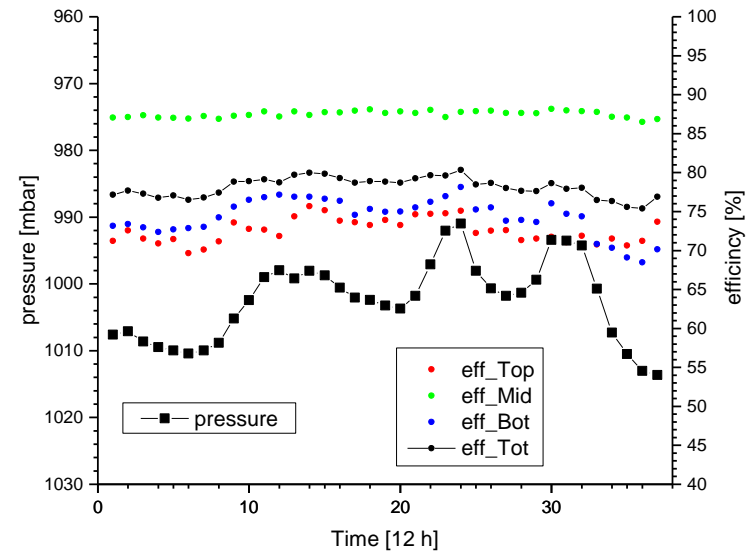
Station C, R169: 37x12h = 444 h of data



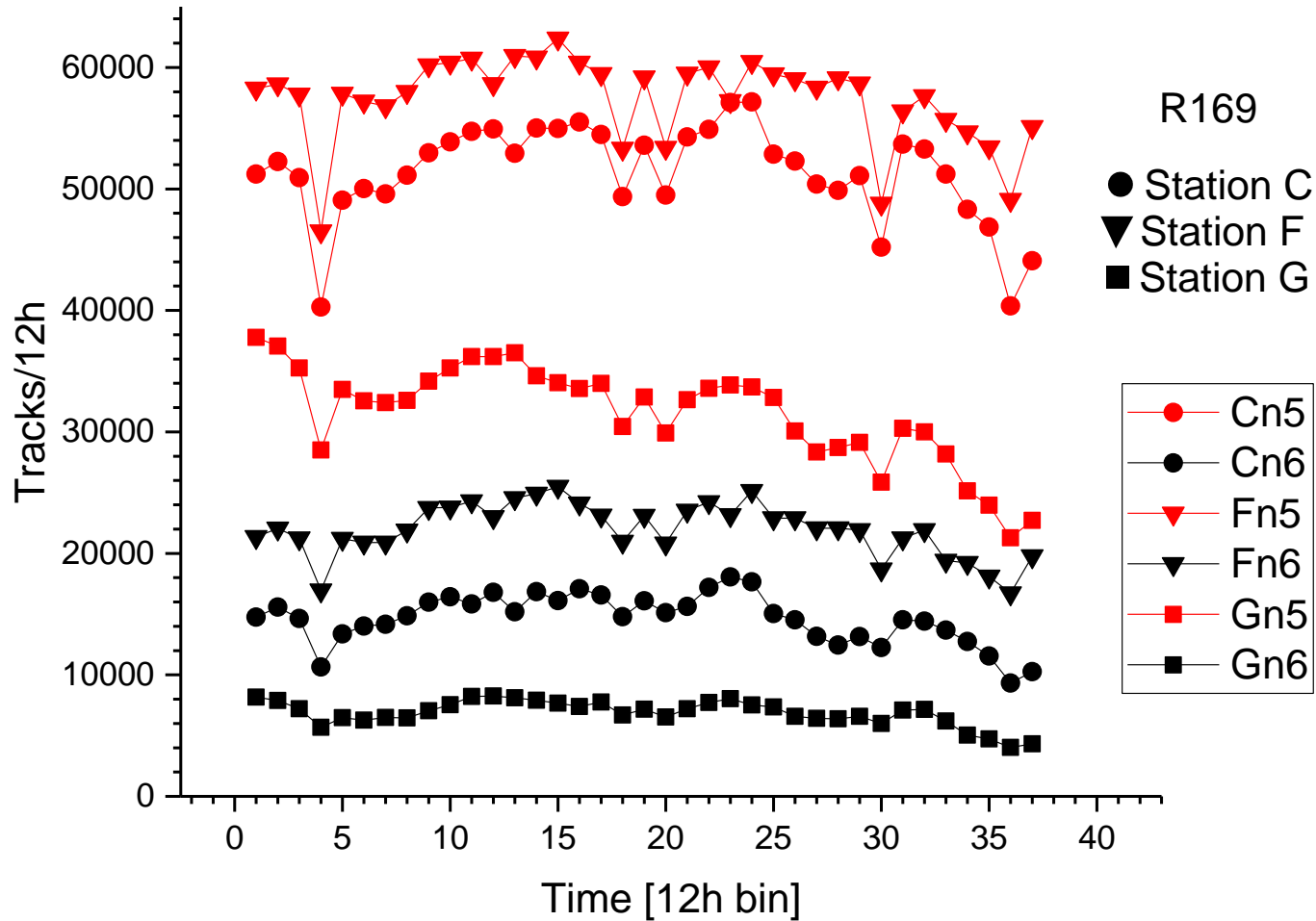
Station G, R169: 37x12h = 444 h of data



Station F, R169: 37x12h = 444 h of data

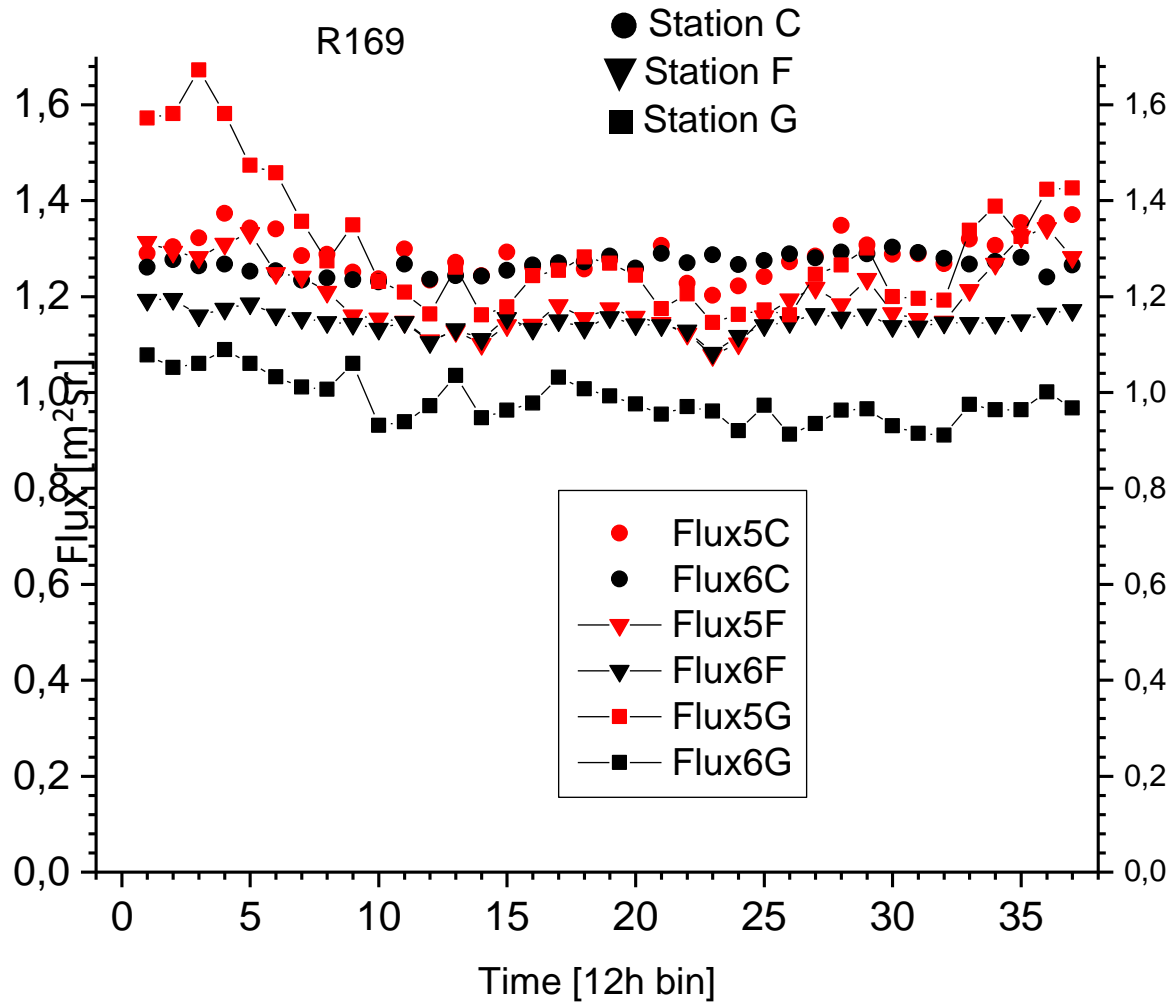


Tracking – single muon rates in three Stations using 5 and 6-hit tracks.



Preliminary

Tracking - single muon flux



Example

- the number of detected multi-muon tracks in run R246 (391h):

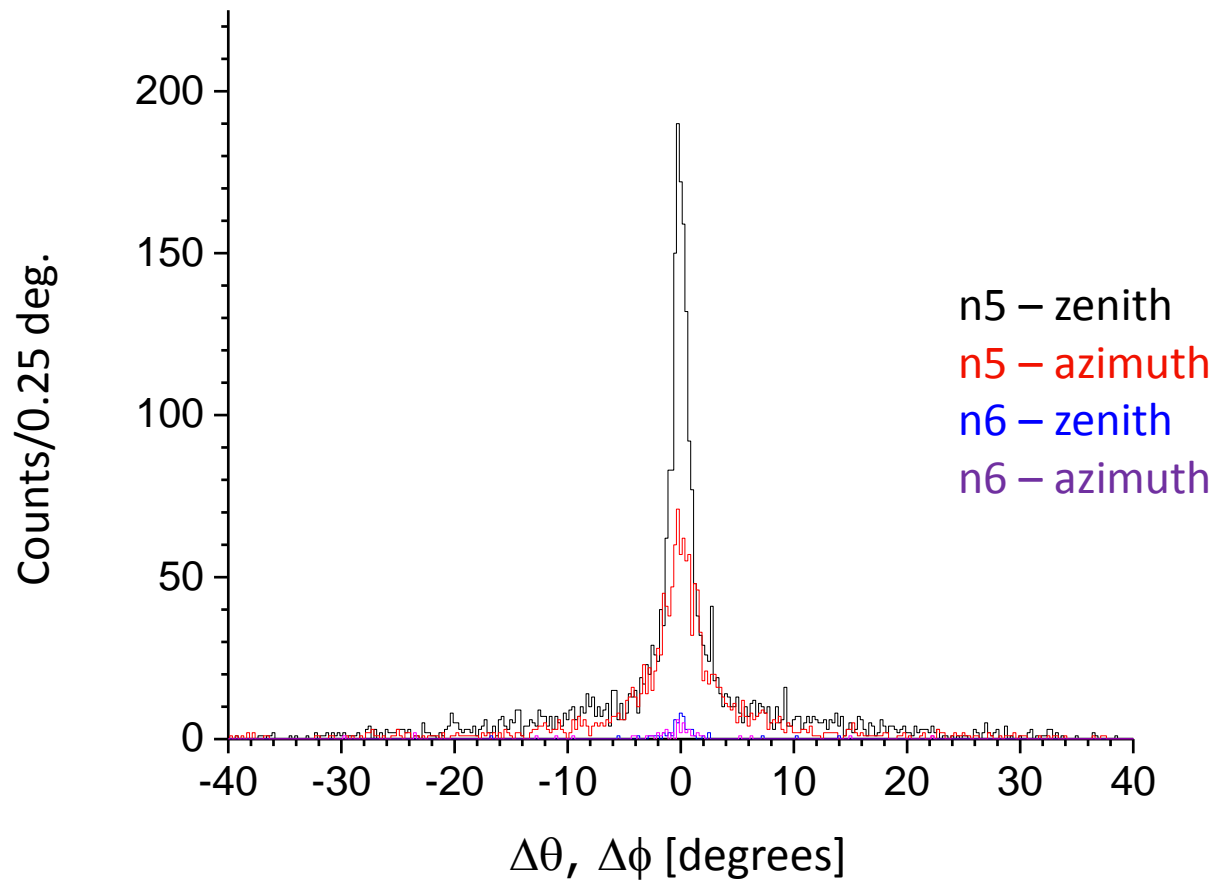
Multiplicities (Mult.) of 5, 6 and 5&6-hit track events in each Station:

	R246-C			R246-F			R246-G		
Mult.	5-hit	6-hit	5&6-hit	5-hit	6-hit	5&6-hit	5-hit	6-hit	5&6-hit
1	1736543	371113	1985682	1578439	346277	1824135	2049245	456026	2360382
2	797	30	61482	819	25	50849	1011	40	73093
3	44	1	225	34	0	197	48	0	281
4	3	0	30	3	0	18	5	0	23
5	0	0	2	1	0	3	2	0	5
6	1	0	2	0	0	2	1	0	3
7	0	0	0	0	0	0	0	0	1
8	0	0	0	0	0	0	0	0	0
9	0	0	0	0	0	0	0	0	0
10	0	0	0	0	0	0	0	0	0

Note: the numbers of detected tracks, no efficiency etc. corrections used.

In the composition analysis measured efficiencies are used as an input in the simulation codes.

Angular distributions $\Delta\theta$, $\Delta\phi$ (zenith and azimuth), Station G, all showers with more than one track.



Hardware:

- LST-electronics in progres,
- Last four detector stations will record data by the end of this year.

Software:

Data analysis can/will be carried out with the small number of program packages:

- + calibrations (two different codes for positions and effs.),
- + tracking,
- + shower simulation.

The next step is to implement SC16s (fully) into the tracking code.

Thank you

Centre for Underground Physics in Pyhäsalmi



EMMA collaboration + SAB, May 2015

Station G, R218: 66x12h = 792h of data

