



EDELWEISS-III Experiment Status and First Low WIMP Mass Results

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Detection of the energy deposited WIMP from galactic halo from galactic halo

v~220 km/s

 $v \sim 0 \text{ km/s}$

due to elastic scattering off target nuclei

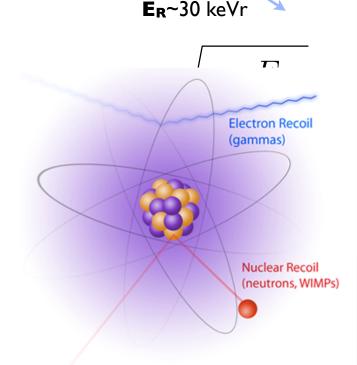
- Elastic scattering of a WIMP deposits small amount of energy into recoiling nucleus (~ few 10s of keV)
- Expected rate:
 < 1 interaction per kg per year
- Radioactive background of most materials gives higher rate

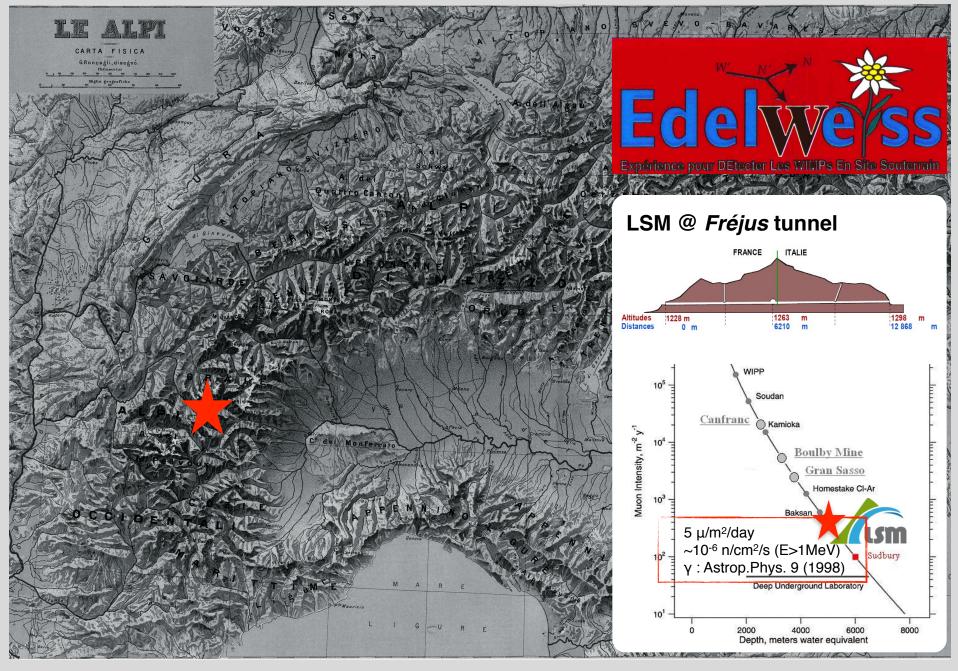




 θ_{R}

WIMP





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Clean room (Rn) with deradonized air supply

with deradonized air supply (from 10 Bq/m³ $\rightarrow \approx$ 30 mBq/m³)

Active muon veto (μ) 97.7% geometric coverage N^{μ -n</sub> = 0.6^{+0.7}_{-0.6} evts (90%CL, 3000kg.d) Astropart. Phys. 44 (2013) 28}

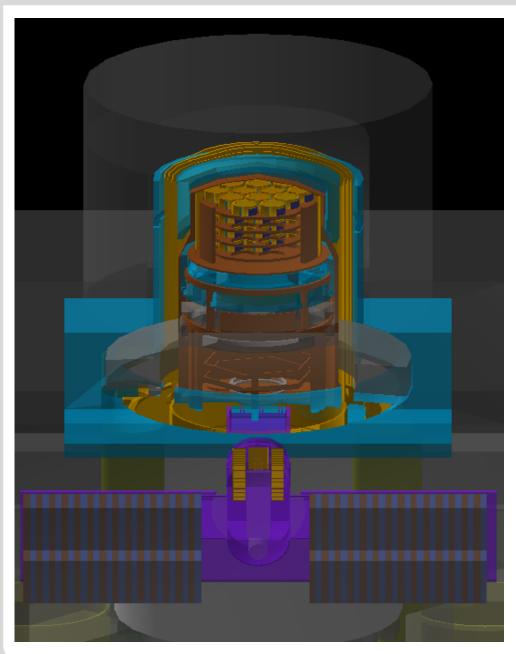
Polyethylene shielding (n) 50cm for moderation

Lead shielding (β, γ) 18cm + 2cm ancient lead

Copper cryostat (β , γ) thermal shielding

- extra 10 cm below detectors PE shield
- extra 15 cm Roman Pb (1K)

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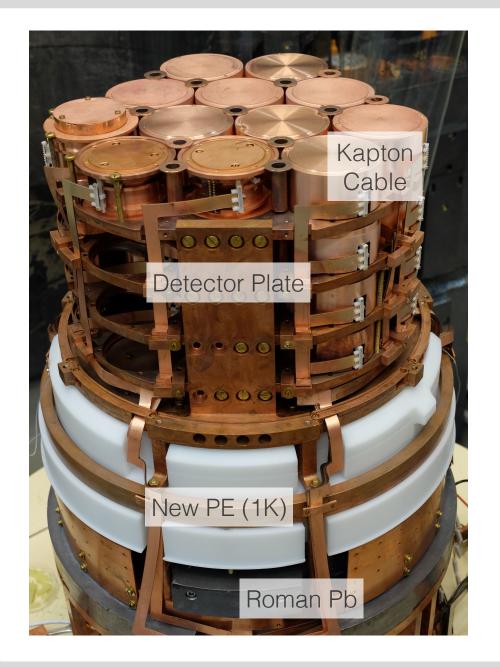
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Detectors within the Cryostat



WIMP search

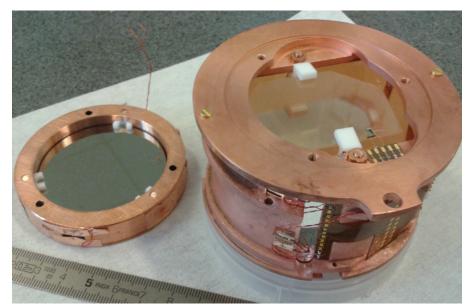
Full Inter-Digitized 800 g HP-Ge Detector



Diameter: 7 cm

$0\nu\beta\beta$ of ¹⁰⁰Mo

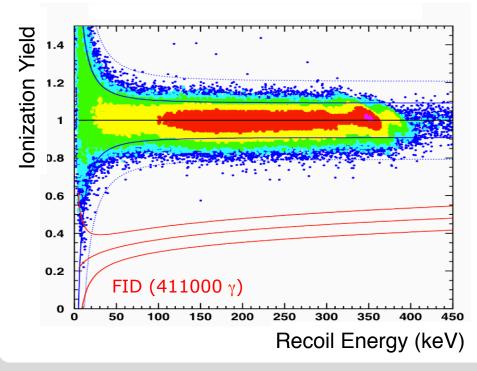
313g ZnMo04 bolometer

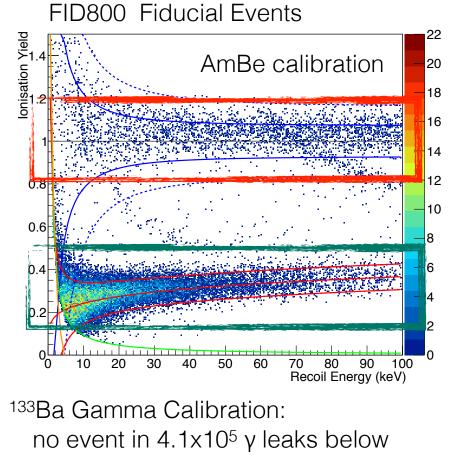


Background Rejection

Most backgrounds (e, γ) produce electron recoils Yield (Ionization/recoil) ~1 WIMPs and neutrons produce nuclear recoils

Yield (Ionization/recoil) ~0.3





-> FID gamma's rejection factor < 6x10⁻⁶

ionization yield of 0.5

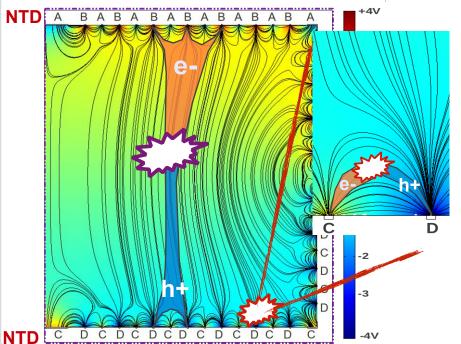
J Low Temp Phys (2012) 167:1056-1062

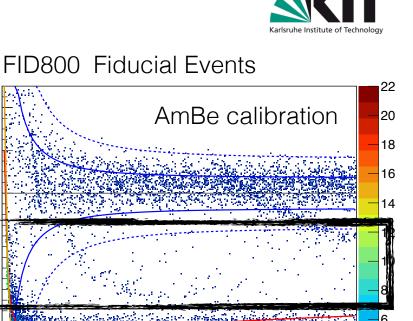
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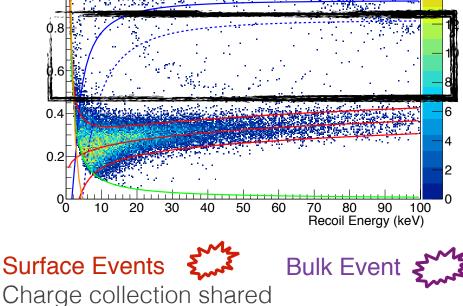
Background Rejection

Particles that interact close to the "surface layer" result in reduced ionization yield.

Surface events can be identified via the ionization signal thanks to ID electrodes







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between one veto and its

electrodes, e.g. C & D

neighbor fiducial

onisation Yield

Charge collected

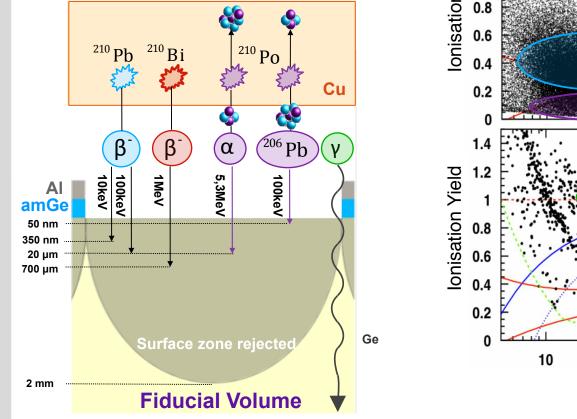
electrodes B & D

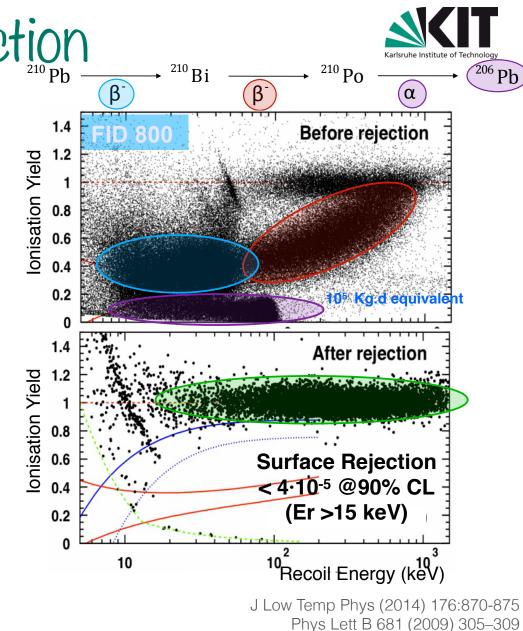
on fiducial

Background Rejection

Installed ²¹⁰Pb implanted on Cu cover facing one detector

Allows performance verification of surface event identification





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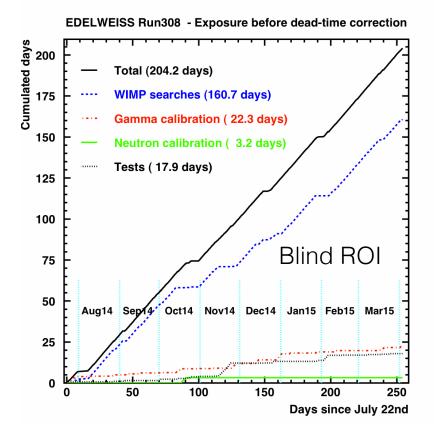
R308 Status



Physics data since summer 2014: 36 x 800 g detectors installed in cryostat 24 x 800 g detectors cabled

Low WIMP mass search:

- Eight months of data taking
- Blind analysis
- Eight detectors with good baselines and low thresholds
- 582 kg·day (fiducial)
- Boosted decision tree (BDT) and a 2D profile likelihood analysis performed



Data Selection

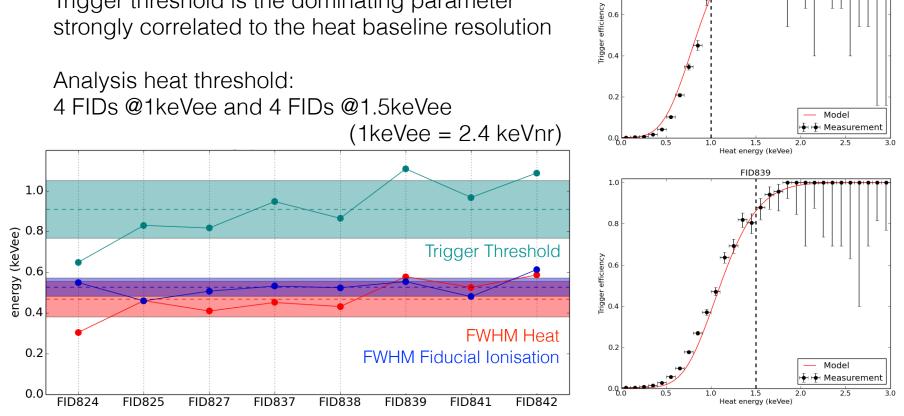


FID825

0.8

- High sensitivity to the lowest WIMP mass achievable
- Good sensitivity up to WIMP mass ~20 GeV
- Dataset as homogenous as possible

Trigger threshold is the dominating parameter strongly correlated to the heat baseline resolution



BDT Analysis

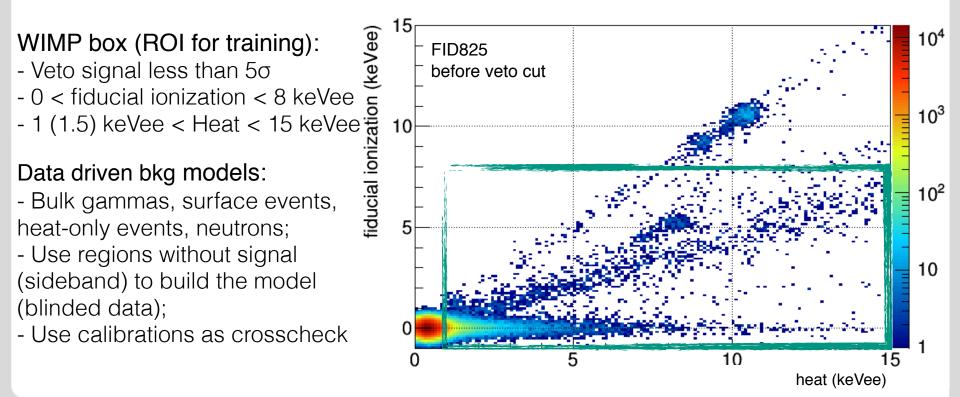
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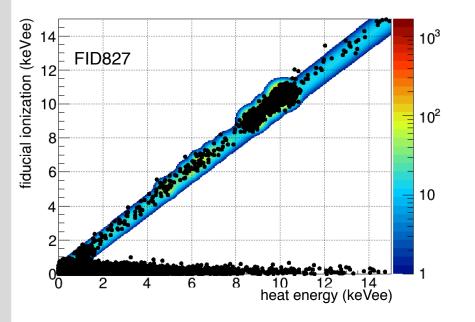
High statistics simulation for BDT training

- Include individual detector effects: trigger, time dependent noise.
- Six variables (4 ionization +1 heat + 1 heat-only event rate) for signal/background discrimination

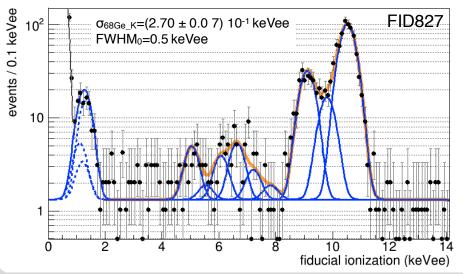
- Model WIMP signal and backgrounds







Lines @ 10.37, 9.66, 8.98, 7.71, 7.11, 6.54, 5.99, 5.46, 4.97 keVee

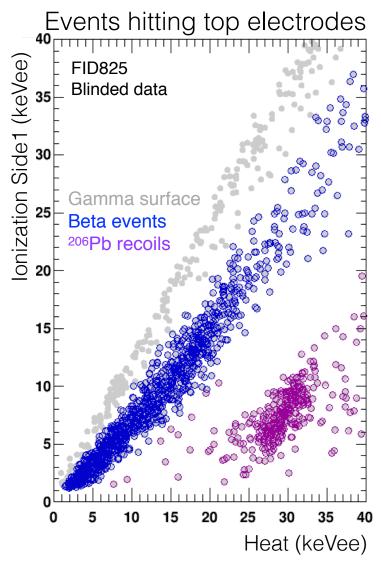


Bulk Gammas

- Fiducial selection
- Fit in [3,15] keVee heat and extrapolation of the flat component down to 0 keV
- L-shell cosmogenic lines with an intensity derived from the K-shell intensity
- The main lines considered are from ⁶⁸Ge, ⁶⁸Ga and ⁶⁵Zn with corresponding L-shell lines at 1.10, 1.19 and 1.30 keV, and a L/K relative intensity of 0.11*

Surface Events



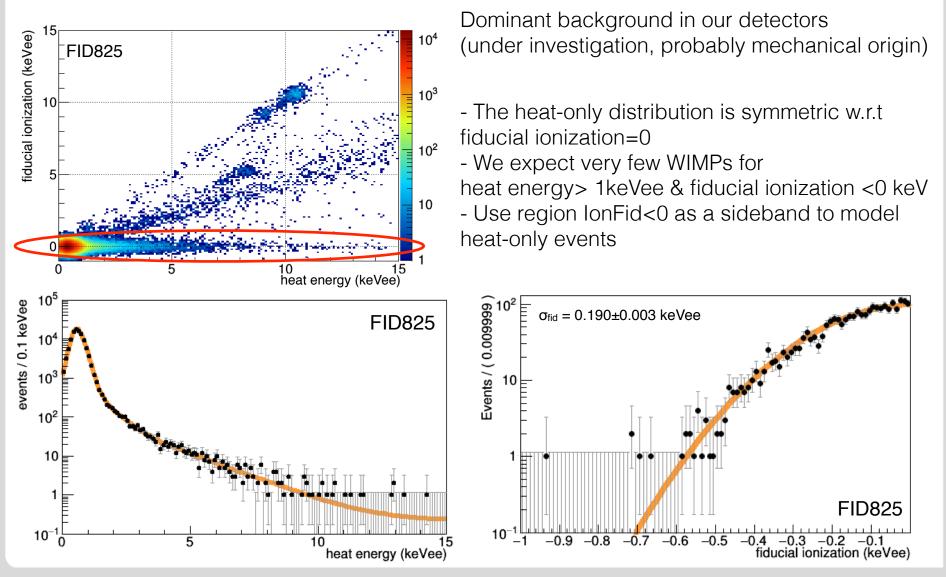


Background models for top and bottom sides:

- **Betas**: spline adjustment in the energy range [4,25] keVee and extrapolation down to 0 keV.
- Lead: adjust the data to a constant plus a gaussian peak in the [10,35] keVee heat energy range and extrapolation to lower energies.
- Surface Gammas: fit in [3,15] keVee heat of a flat component + lines. The intensity of lines is scaled from intensity of corresponding fiducial lines w/ fiducial mass fraction.

Heat-Only Events

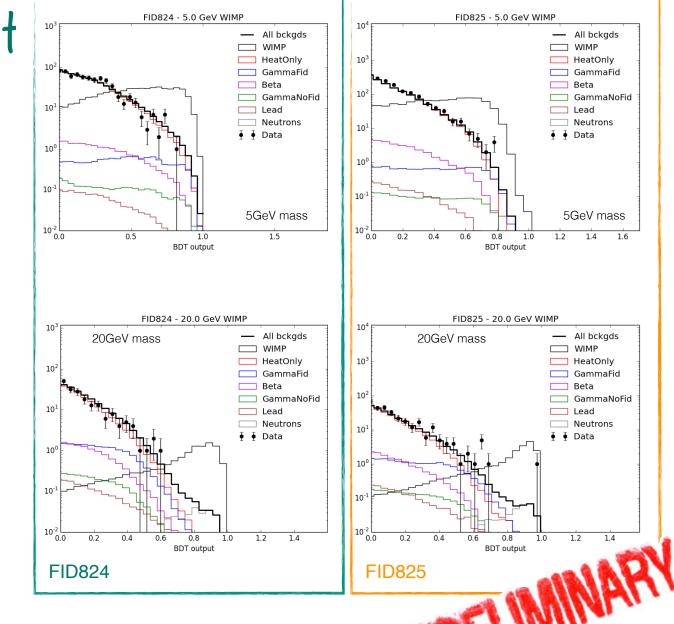




BDT Output

One BDT distribution per WIMP mass of 4, 5, 6, 7, 10, 15, 20 and 30.0 GeV.

Backgrounds are normalized to the expected number of events for that given detector and data selection. WIMP signal not normalized.



BDT Output

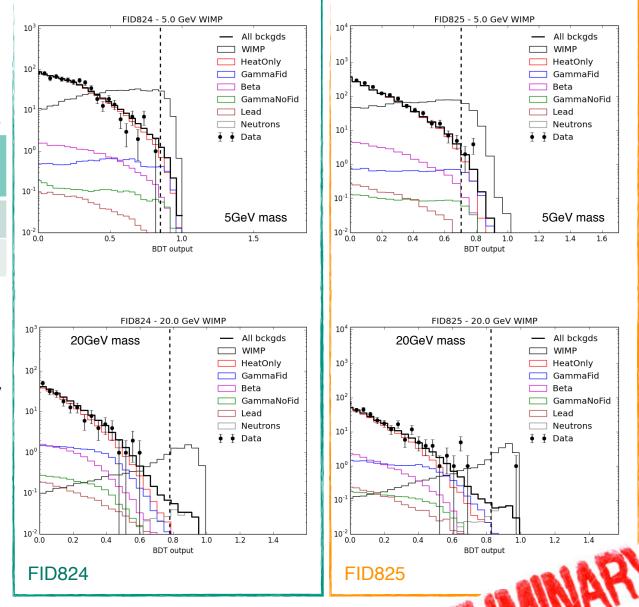
After BDT and WIMP box cut in 8 FIDs

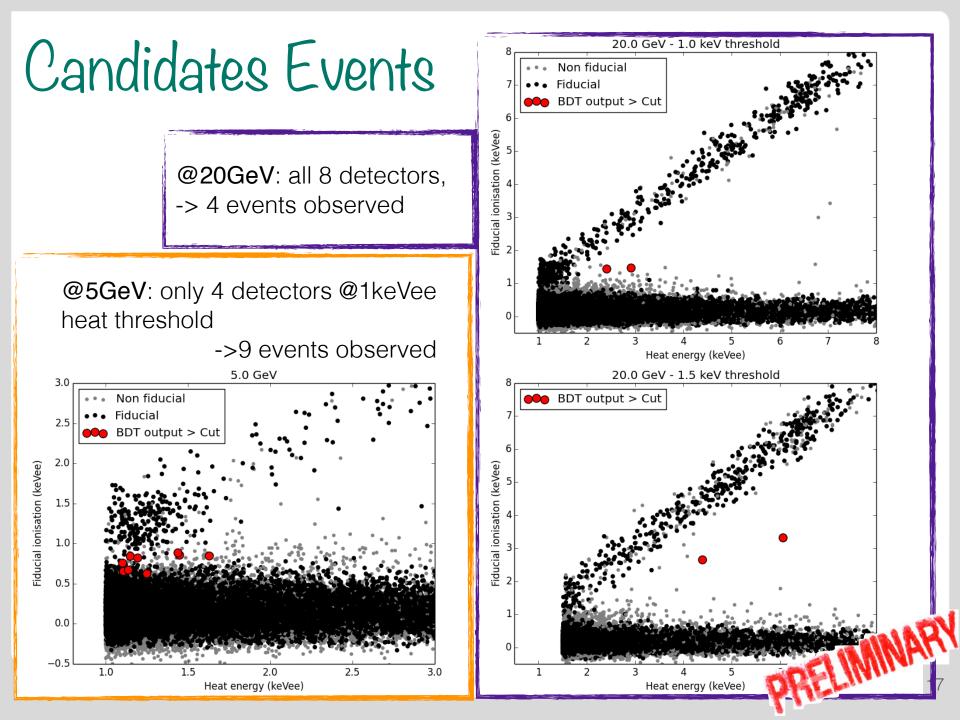
	N_bkg expected	N_bkg observed	p_value (stat only)
5 _{GeV}	6.14	9	0.17
20 _{GeV}	1.35	4	0.10

Dominant background:

Low WIMP mass: heat-only events and cosmogenic gamma lines

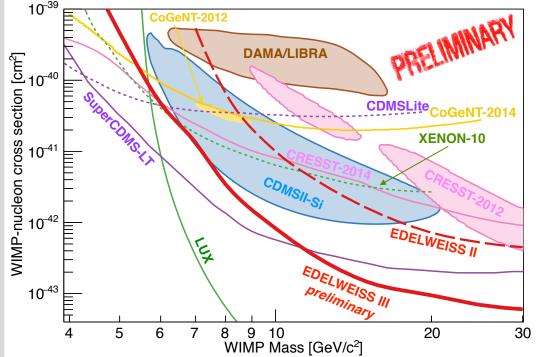
High WIMP mass: radiogenic neutrons (preliminary systematic of 45%)





Low Mass WIMP Limit





- Preliminary limit
- w/o background subtraction
- Poisson limit

The preliminary 90%CL limit achieved for spin-independent WIMP-nucleon varies from 4.6×10^{-40} cm² at 5 GeV/c² to 6.2×10^{-44} cm² at 30 GeV/c².

Cross checks with a 2D profile likelihood analysis is ongoing and already shown good agreement

Summary & Outlook R308



- Low energy WIMP mass (4-30GeV) analysis performed for 582 kgd (fiducial).
 - A factor 40 (@5GeV) and 8 (@30GeV), w.r.t to the previous generation of ID detectors with an exposure of 113 kgd (fiducial).
 - Crosscheck w/ a 2D profile likelihood analysis is ongoing.
 - Post-unblinding checks are ongoing.
- High energy WIMP mass analysis results soon.

Current Run (R309)

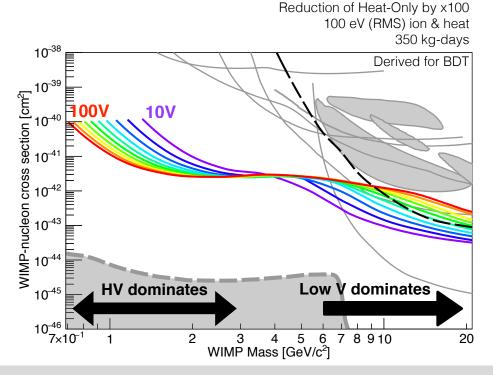
DAQ resumed in June 2015:

23 FID800 (12 new) One FID200 for 'High-Voltage' R&D, i.e. Neganov-Luke amplification

HEMT R&D to lower ionization threshold down to $\sigma_{ion} = 100 \text{ eV}$

R&D on heat sensor goal: σ_{heat} = 100 eV and HV (Luke-Neganov) to reduce recoil threshold

R&D to reduce heat-only events



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CEA Saclay (IRFU & IRAMIS) CSNSM Orsay (CNRS/IN2P3 & Paris Sud) IPNL Lyon (CNRS/IN2P3 & Univ. Lyon 1) Néel Grenoble (CNRS/INP) LPN Marcoussis (CNRS)



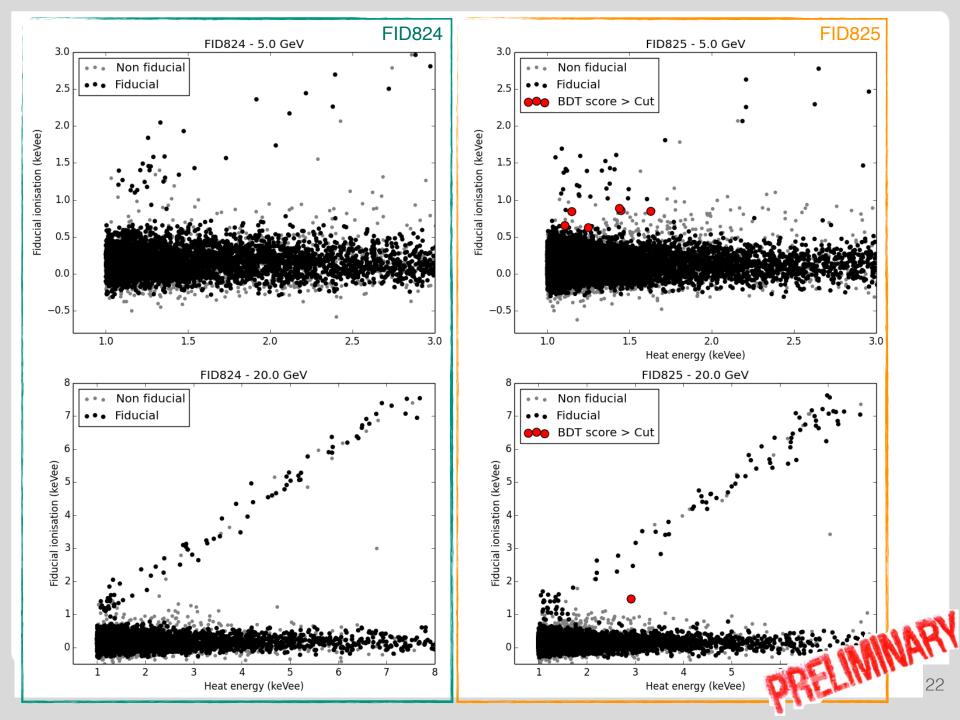
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Back-up

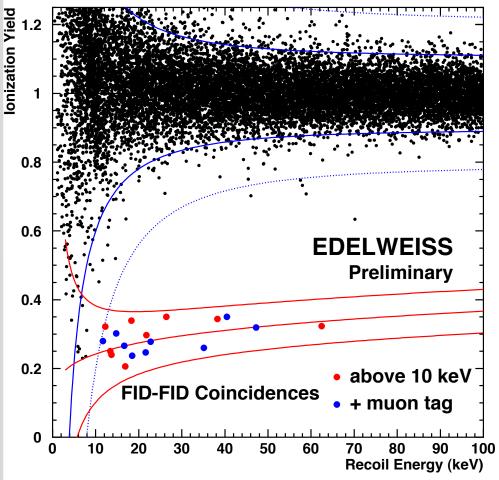
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Neutrons





- During the WIMP search, we see 9 multiple nuclear recoil events (excluded from the search data set) after muon cut, in 17 detectors in 1300 kgdays

 It has been used as normalization factor in simulation, BDT training and BDT cut optimization along with the single-to-multiple ratio

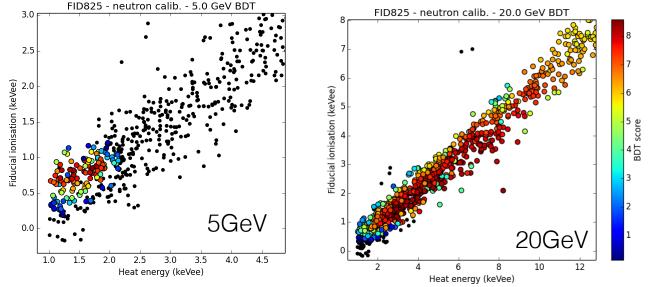
 Single-to-multiple ratio from radiogenic neutron simulations varies between
 FIDs. An average has been considered

Systematics : sqrt(9) + large variation is single-to-multiple ratio





AmBe neutron calibration and BDT output



We can see NR at low energies with a good efficiency in the low mass WIMP analysis

Low WIMP mass: neutron is a negligible background. Events passing the BDT cut are below 2keVee heat energy and we would expect radiogenic neutrons at higher energy

High WIMP mass: BDT cut at ~7, dark red events are passing the cut.

