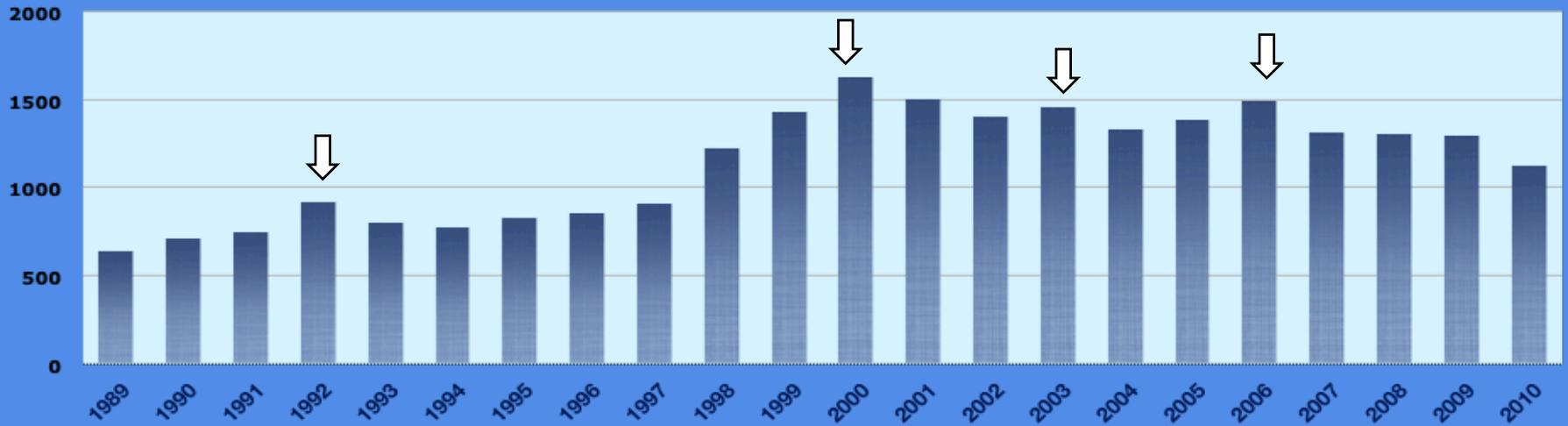


TAUP 2011  
Munich, Germany

# Neutrinos: Phenomenology and Interpretation

Eligio Lisi  
INFN, Bari, Italy

# N. of neutrino preprints since first TAUP (1989) - from INSPIRE

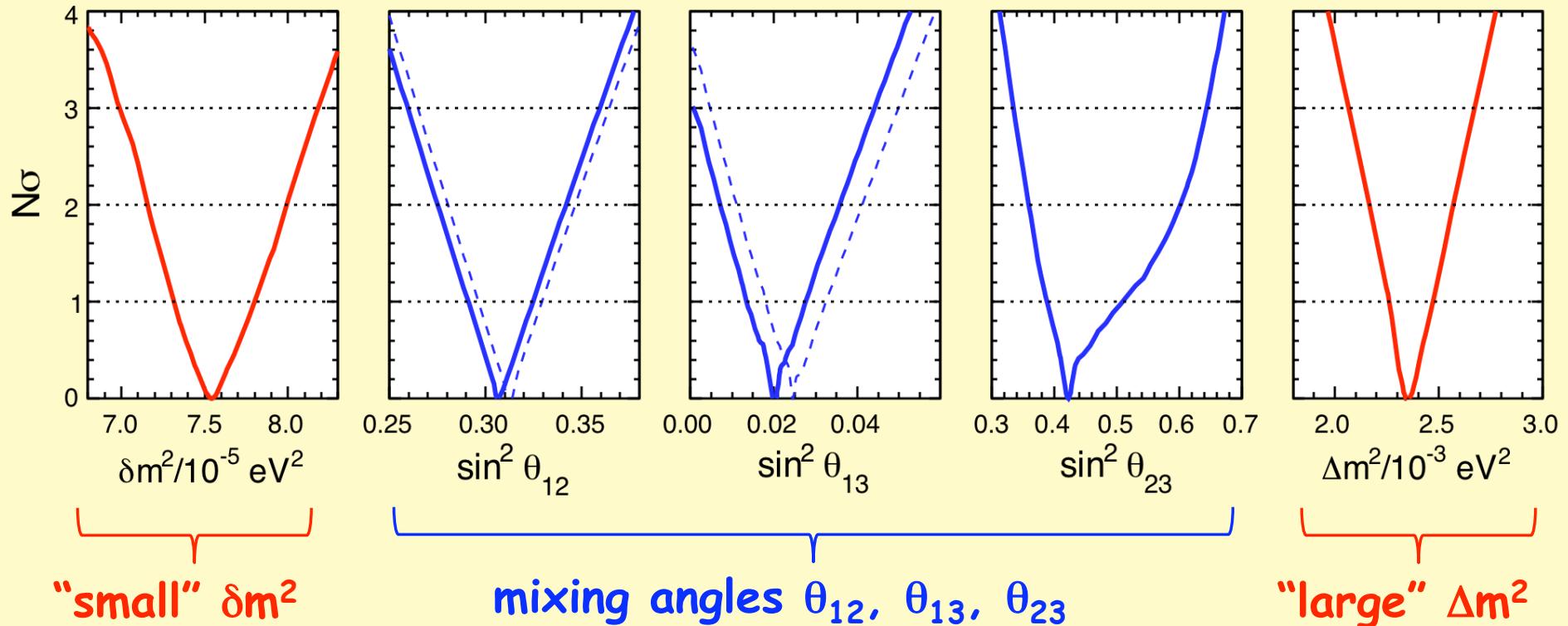


2011 likely to become another "peak year:"  $\theta_{13}$ , sterile  $\nu$  hints, ...

→ topics of this talk:  $3\nu$  &  $3\nu+\nu_s$

# Status of $3\nu$ oscillation parameters, circa 2011

Results of recent global data analysis in terms of  $N\sigma = \sqrt{\Delta\chi^2}$   
 (the more linear and symmetric, the more gaussian errors)



G.L.Fogli, E.L., A.Marrone, A.Rotunno (Univ. & INFN, Bari), A. Palazzo (TUM, Munich)  
 "Evidence of  $\theta_{13} > 0$  from global neutrino data analysis", arXiv:1106.6028v2

## Numerical $1\sigma$ , $2\sigma$ , $3\sigma$ ranges:

Parameter	$\delta m^2/10^{-5} \text{ eV}^2$	$\sin^2 \theta_{12}$	$\sin^2 \theta_{13}$	$\sin^2 \theta_{23}$	$\Delta m^2/10^{-3} \text{ eV}^2$
Best fit	7.58	0.306 (0.312)	0.021 (0.025)	0.42	2.35
$1\sigma$ range	7.32 – 7.80	0.291 – 0.324 (0.296 – 0.329)	0.013 – 0.028 (0.018 – 0.032)	0.39 – 0.50	2.26 – 2.47
$2\sigma$ range	7.16 – 7.99	0.275 – 0.342 (0.280 – 0.347)	0.008 – 0.036 (0.012 – 0.041)	0.36 – 0.60	2.17 – 2.57
$3\sigma$ range	6.99 – 8.18	0.259 – 0.359 (0.265 – 0.364)	0.001 – 0.044 (0.005 – 0.050)	0.34 – 0.64	2.06 – 2.67

Typical  $1\sigma$  accuracy [defined as 1/6 of  $\pm 3\sigma$  range]

$\delta m^2$	$\sin^2 \theta_{12}$	$\sin^2 \theta_{13}$	$\sin^2 \theta_{23}$	$\Delta m^2$
2.6%	5.4%	$0.7 \times 10^{-2}$ (abs.err.)	12%	4.3%

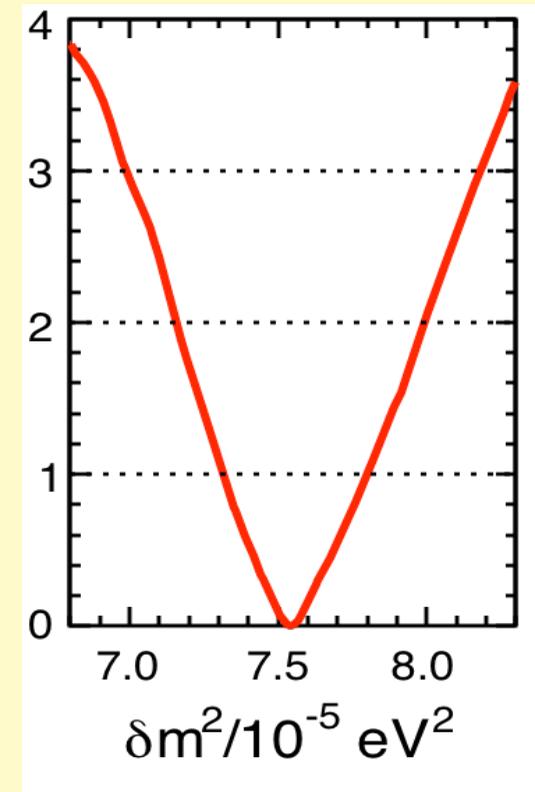
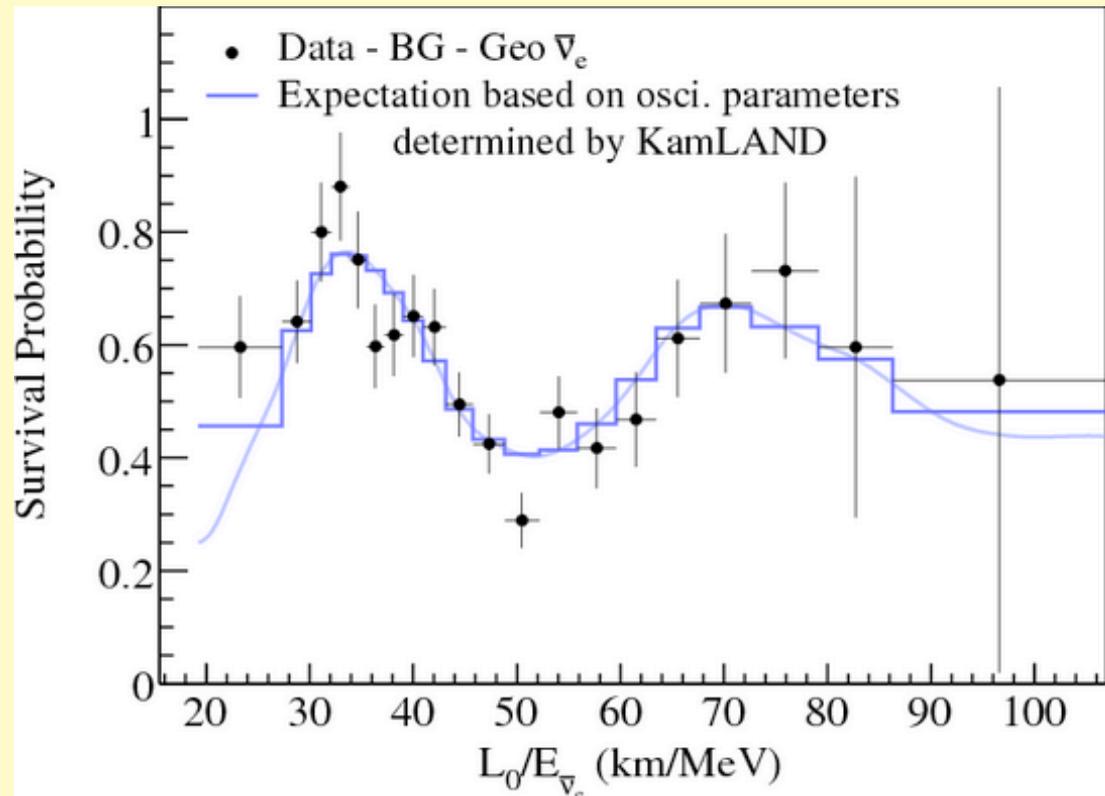
We are already in the precision era for  $\nu$  physics!

More details on  $3\nu$  parameters  $\rightarrow$

# Zooming in on $\delta m^2 = \Delta m^2_{12}$

Vacuum oscillation frequency  
 $\delta m^2/2E$ : dominated by KamLAND  
 (interesting by-product: geoneutrinos!)

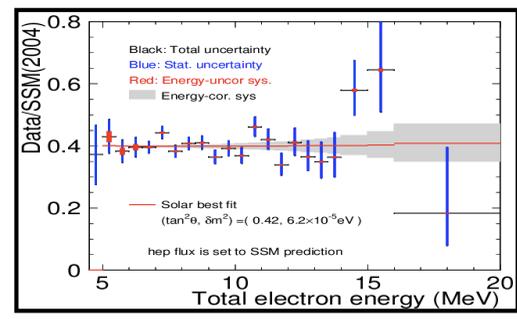
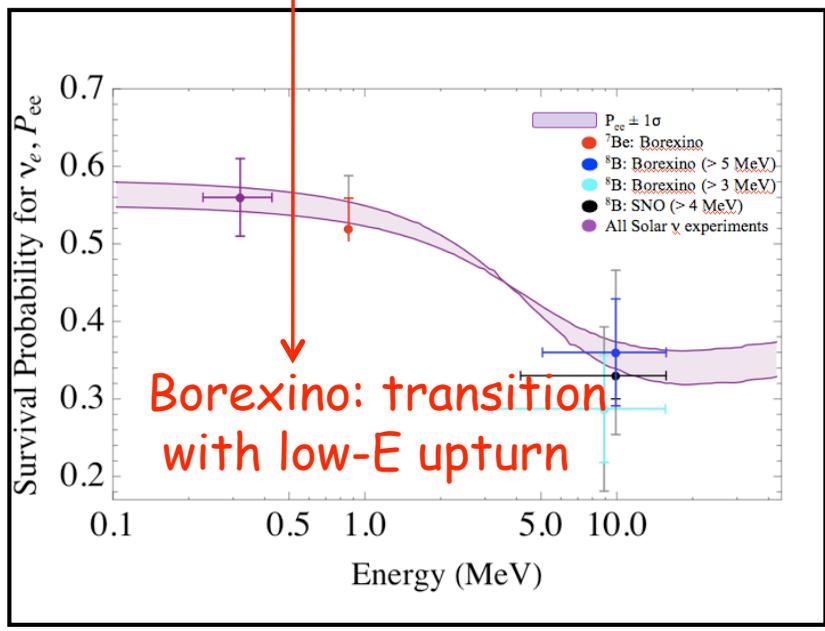
Accuracy:  
 near ultimate limit  
 ( $\sim 2\%$  energy scale)



Vacuum frequency  $\delta m^2/2E$  (unknown sign) interferes with  $\nu_e$  interaction energy  $\sqrt{2}G_F N_e$  (known sign) in solar matter: MSW

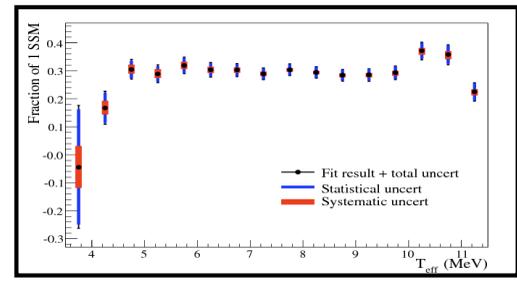
→ Get  $(\nu_1, \nu_2)$  hierarchy: state with largest  $\nu_e$  component is the lightest of the two - conventionally,  $\nu_1$

### Main features of matter effects: established Spectral details: need higher stat, lower E



↓  
**SK:**  
 ~flat

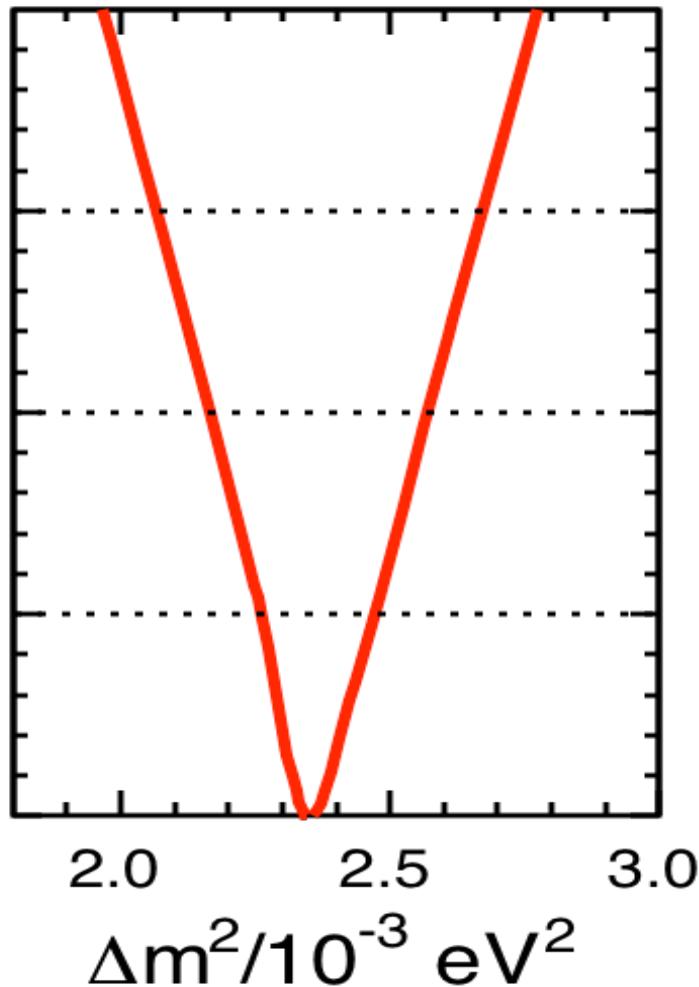
↓  
**SNO:**  
 ~downturn



Here at TAUP 2011: Final SNO data analysis and first Borexino pep flux data!

# Zooming in on $\Delta m^2 \approx |m^2_3 - m^2_{1,2}|$

$\Delta m^2$  (ATM+K2K+MINOS)



$$\sigma(\Delta m^2) \sim 1.3 \cdot \delta m^2 !$$

Setting  $\delta m^2 = 0$  is no longer a reasonable approximation in fitting  $\Delta m^2$ .

For  $\delta m^2 > 0$  need to declare  $\Delta m^2$  convention. We use:

$$\Delta m^2 = (\Delta m^2_{31} + \Delta m^2_{32})/2$$

in both normal and inv. hierarchy.

[Changing hierarchy just flips sign of  $\pm \Delta m^2$  with no significant difference in current fit; results marginalized over  $\text{sign}(\Delta m^2)$ .]

Can the “ $\delta m^2$  success story” be repeated for  $\text{sign}(\pm\Delta m^2)$  ?  
 Is the state with smallest  $\nu_e$  component ( $\nu_3$ ) lightest or heaviest?

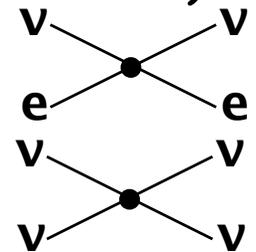
The hierarchy, namely,  $\text{sign}(\pm\Delta m^2)$ , can be probed (in principle), via interference of  $\Delta m^2$ -driven oscillations with some other Q-driven oscillations, where Q is a quantity with known sign.

At present, the only known possibilities (barring new physics) are:

Q =  $\delta m^2$  (e.g., high-precision oscillometry in vacuum)

Q = **Electron density** (e.g., matter effects in Earth)

Q = **Neutrino density** (SN  $\nu$ - $\nu$  interaction effects)



Each of them is very challenging, for rather different reasons.

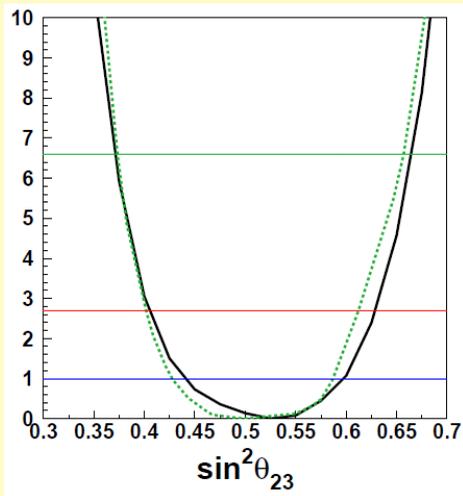
The latter possibility has recently raised increasing interest, being associated with **highly nonlinear flavor evolution effects**.

All of them worth revisiting in the light of possible “large”  $\theta_{13}$  !

# Zooming in on $\sin^2\theta_{23}$ : maximal $\nu_\mu$ - $\nu_\tau$ mixing, or not ?

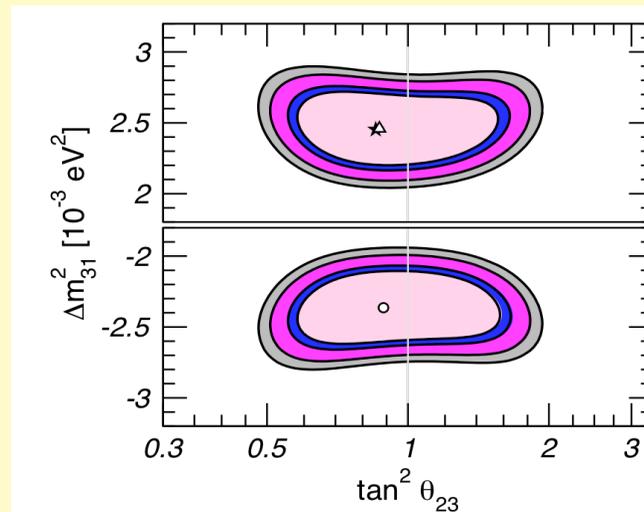
3 $\nu$  analyses including SK-I+II+III data with both  $\theta_{13}\neq 0$  and  $\delta m^2\neq 0$ :

SK@Neutrino2010



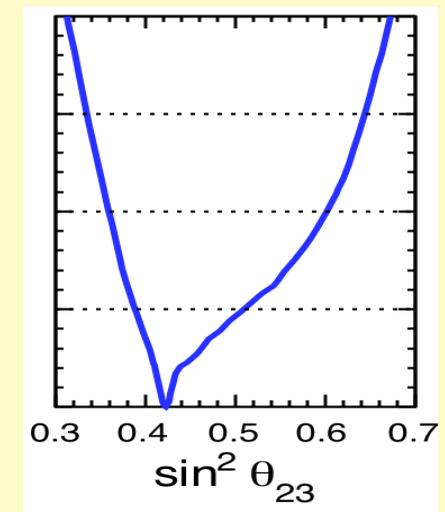
Nearly maximal...

Gonzalez-Garcia et al. 2010



Slightly nonmaximal...

Our analysis, 2011

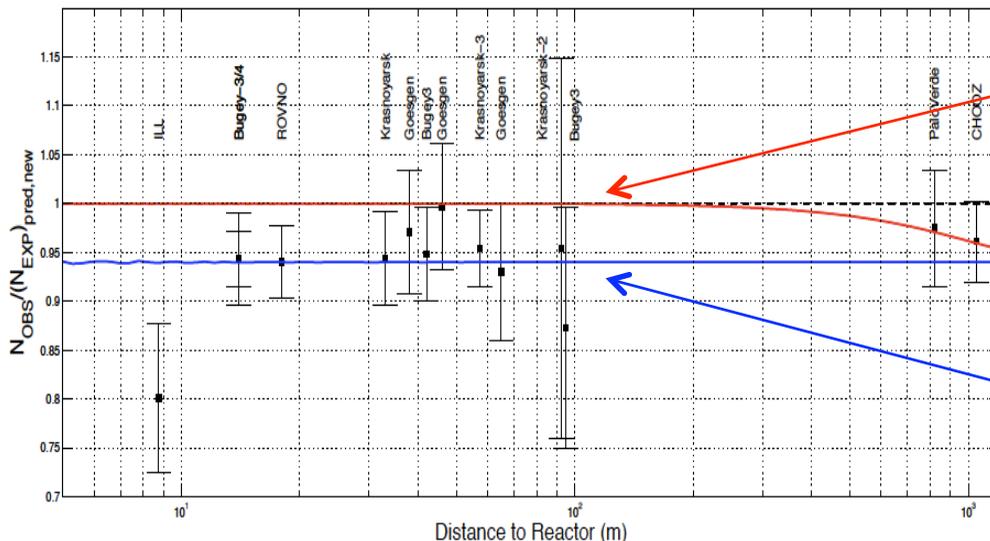


Slightly more nonmaximal...

In general, two "signs" appear to be "fragile" in current best fits:  
 $\text{sign}(\pm\Delta m^2) \rightarrow$  hierarchy,  $\text{sign}(\sin^2\theta_{23}-1/2) \rightarrow$  octant,  
 their effects being comparable to stat/syst errors (approximations ?)

## Zooming in on $\sin^2\theta_{12}$ and $\sin^2\theta_{13}$ : $\nu_e$ mixing

Here, short and long baseline reactor data play an important role.  
**Reactor neutrino flux normalization revision in 2011\***: about +3.5%.  
 Systematic problem in comparing old  $\nu$ -rate and  $\beta$ -decay data:



"New" or "high" normalization agrees with ILL  $\beta$  spectra but not with very-SBL reactor rates

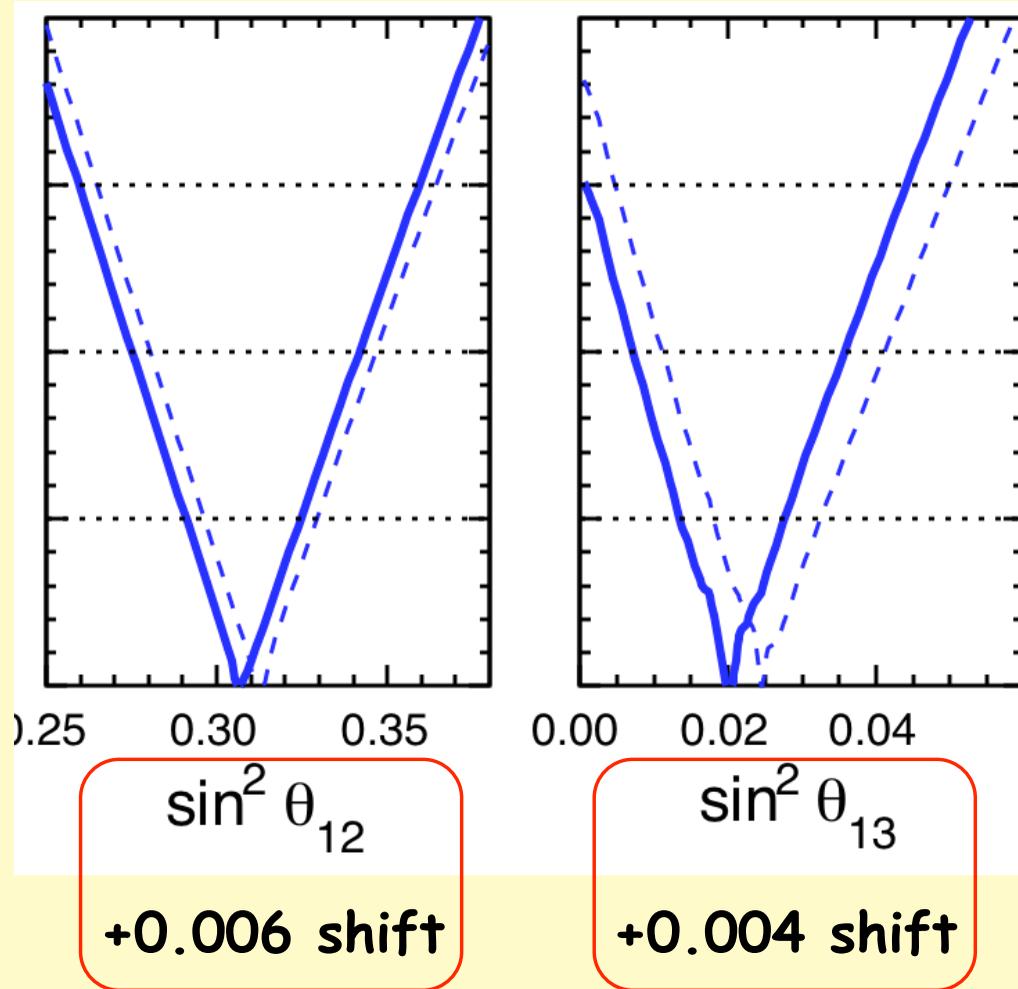
"Old" or "low" normalization agrees with VSBL reactors but not with ILL  $\beta$  spectra

Tension at 2-3 $\sigma$  level.  $\nu_s$  ?

Within  $3\nu$  (no new physics)  $\rightarrow$  ununderstood theo/exp systematics ?  
 May gauge the impact by comparing effects of old/new fluxes  
 Expect: higher flux  $\rightarrow$  more disappear.  $\rightarrow$  larger  $\theta_{12}$ ,  $\theta_{13}$  angles

\*Mention et al., arXiv:1101.2633; Huber, arXiv:1106.0687

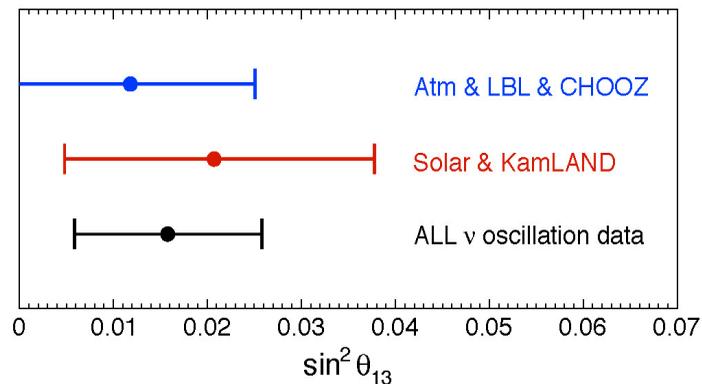
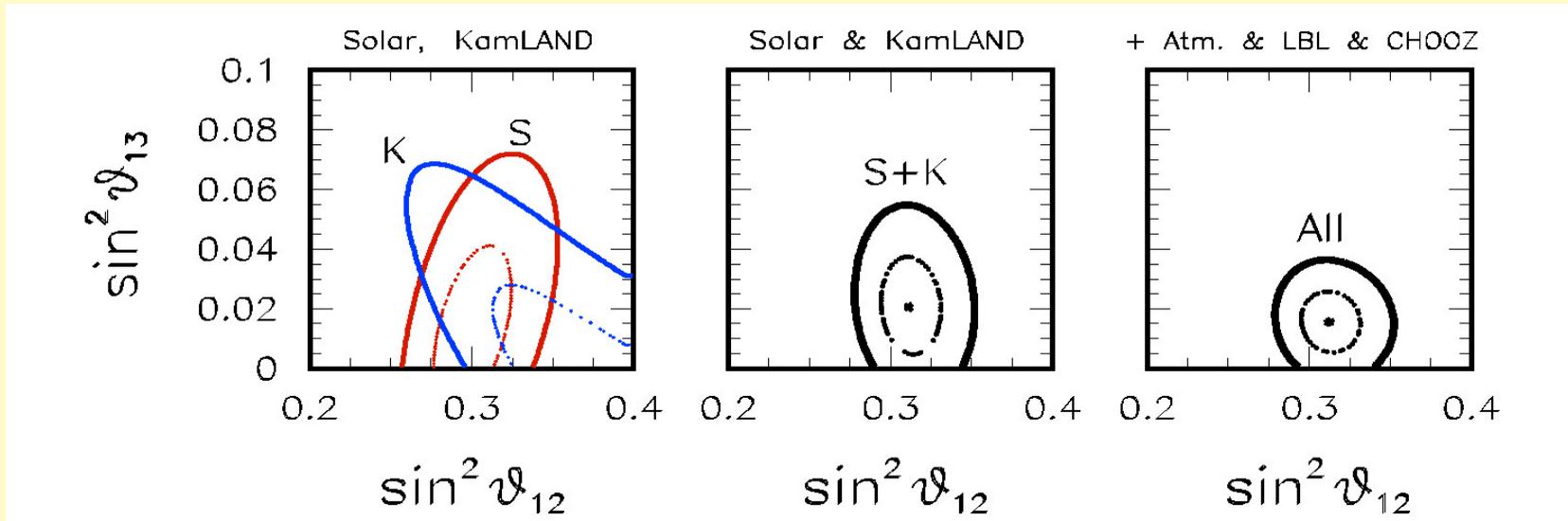
## Passing from "old" to "new" fluxes:



Note: Analyses that privilege "old" or "low" normalization by fitting VSBL reactor data [Schwetz et al., arXiv:1108.1376] tend to prefer lower  $\theta_{13}$

# The increasing convergence of hints for $\theta_{13} > 0$

2008: "Hints of  $\theta_{13} > 0$  from global neutrino data analysis"  
 [Fogli, EL, Marrone, Palazzo, Rotunno, arXiv:0806.2649]



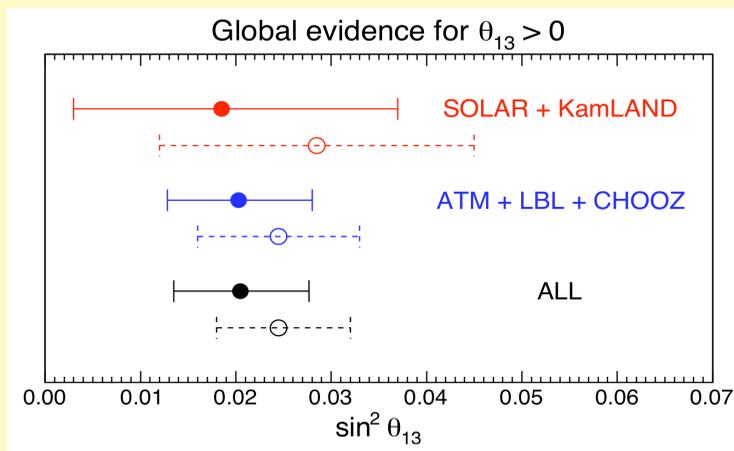
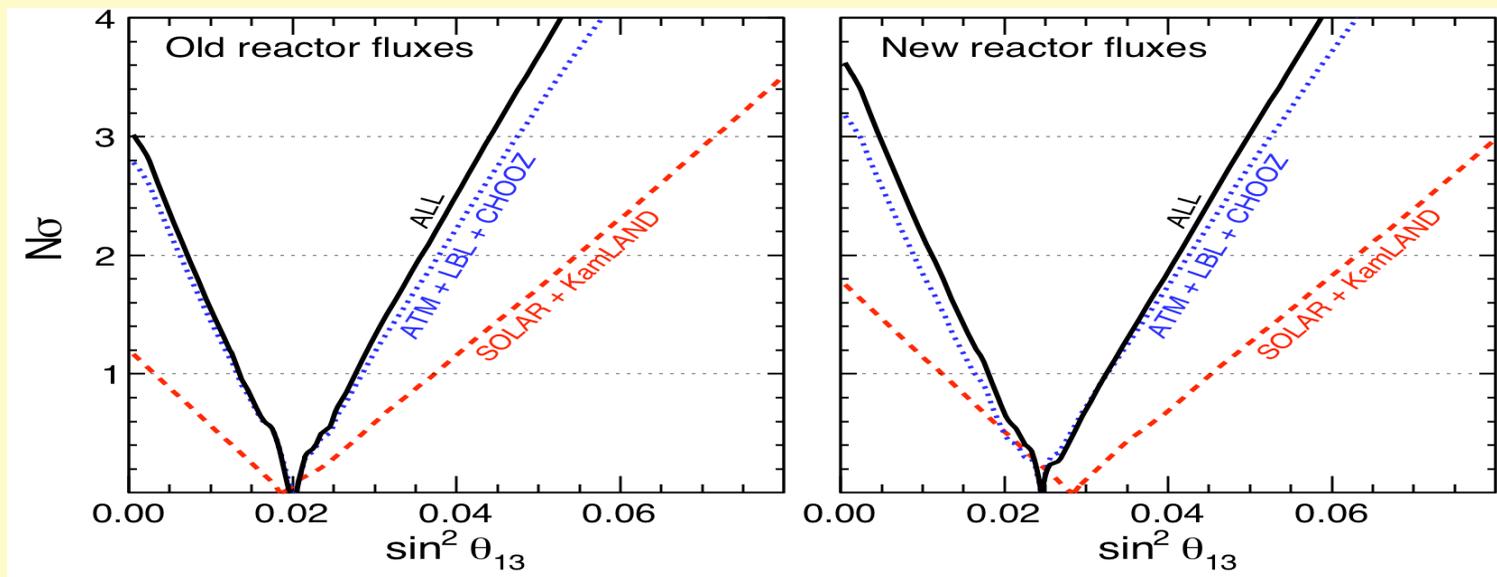
Hint significance: 90% CL.

$$\sin^2 \theta_{13} = 0.016 \pm 0.010$$

[Large literature and debate 2008-2011]

# June 2011 breakthrough: new T2K and MINOS appearance results!

→ "Evidence of  $\theta_{13} > 0$  from global neutrino data analysis"  
 [Fogli, EL, Marrone, Palazzo, Rotunno, arXiv:1106.6028]



Statist. significance  $> 3\sigma$

$$\sin^2\theta_{13} =$$

$0.021 \pm 0.007$  (old fluxes)  
 $0.025 \pm 0.007$  (new fluxes)

## Words of caution, from someone we all miss:



**“Half of all  $3\sigma$  results are wrong.”**

J.N. Bahcall

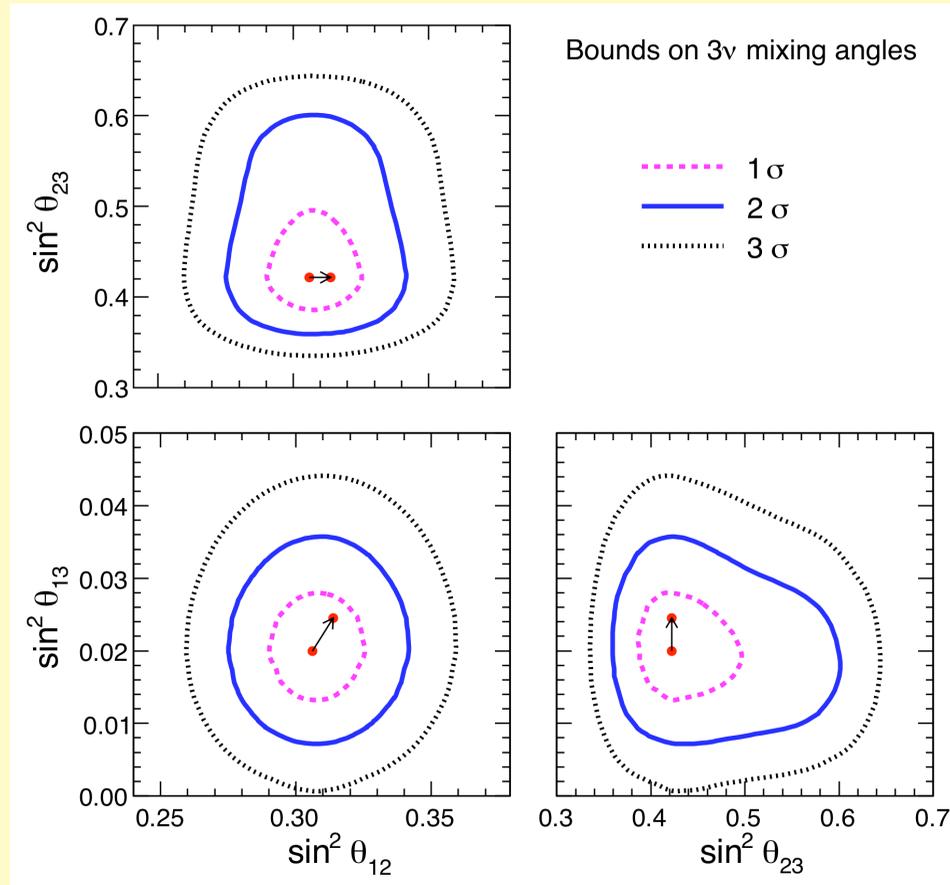
**No doubt: we need further experimental tests of  $\theta_{13} > 0$   
and, in particular, at least a single result at  $>3\sigma$  level !**

But, it would be very surprising if the current evidence were just due to a conspiracy of fluctuations in rather different categories of neutrino oscillation experiments.

In planning future  $\nu$  beams/detectors, useful to consider the likely possibility that  **$\sin^2\theta_{13} \sim 0.02$**

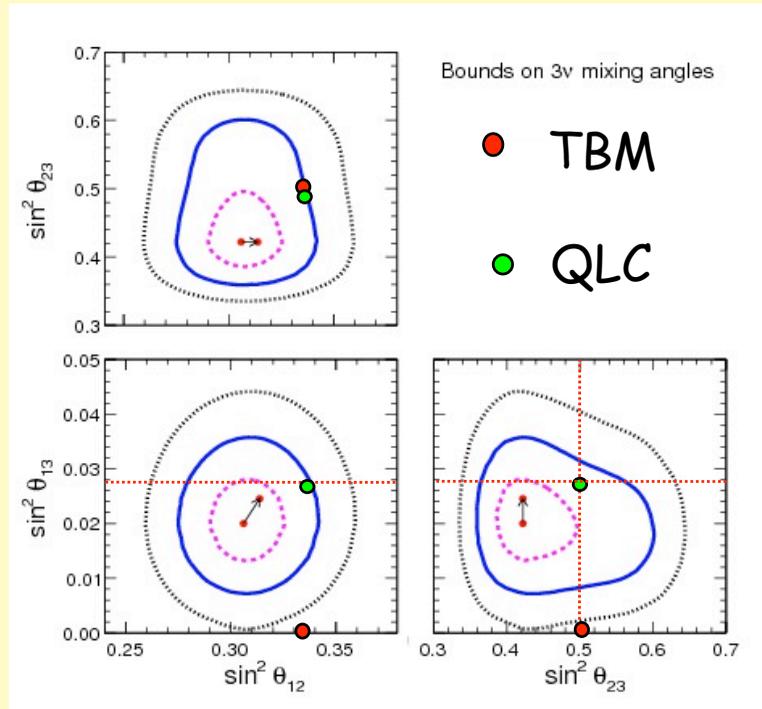
# Overview of 3ν mixing: joint bounds on $\sin^2\theta_{ij}$

(arrows: shifts from "old" to "new" fluxes)

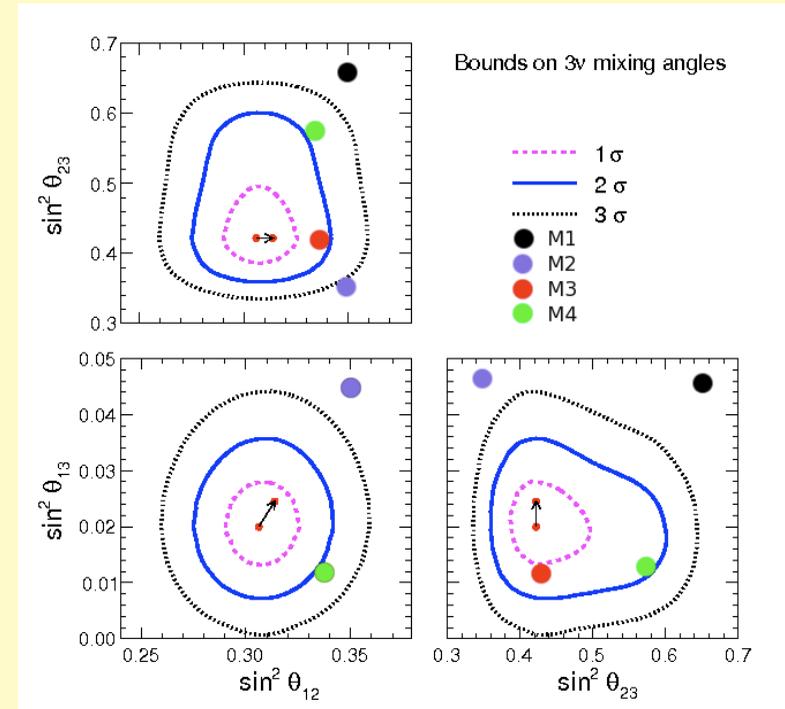


Note slight **anticorrelation** between  $\sin^2\theta_{23}$  and  $\sin^2\theta_{13}$ , induced by T2K+MINOS appearance data which constrain their product.

Several theoretical models undergoing more sever tests, e.g.,



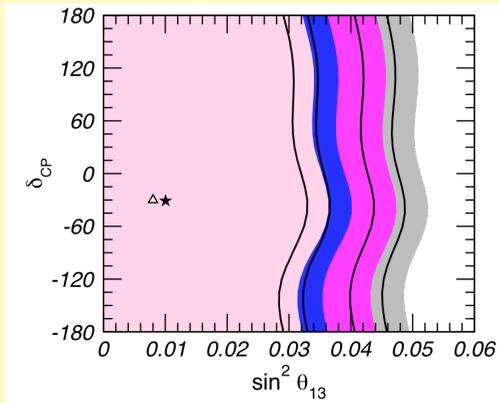
A. Smirnov @ Nufact 2011



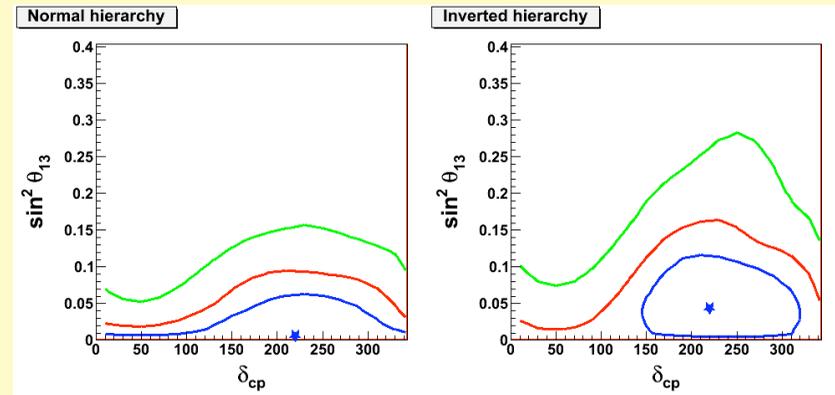
Feruglio et al., arXiv:1107.3486

Many ideas being proposed to make sense (if any) of "large"  $\theta_{13}$ .  
 Can it be linked to some other small parameters, such as  $\theta_{\text{Cabibbo}}$ ?

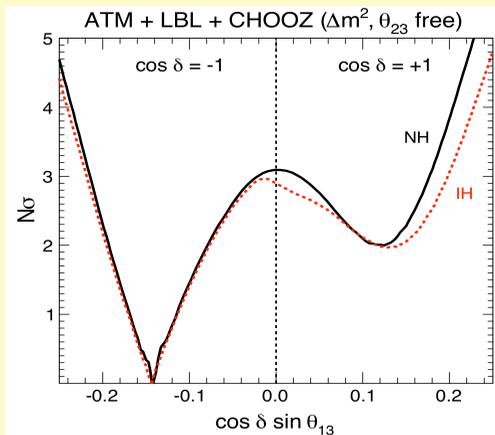
$\theta_{13} > 0$  opens the door to leptonic CP violation ( $\delta$ ).  
 Wild dream: peeking through the door... now?



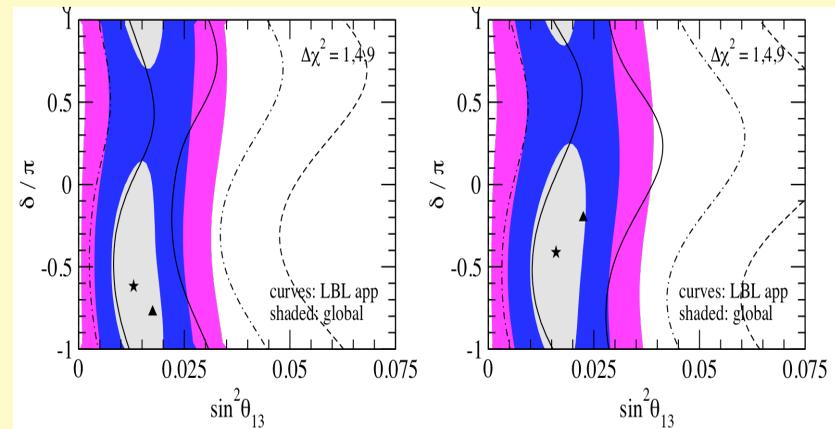
Gonzalez-Garcia et al. 2010



Takeuchi [SK atmospheric] @ Neutrino 2010



Fogli et al., 2011



Schwetz et al., 2011

Emerging sensitivity to  $\delta_{CP}$ ? Time will tell...

# More dreams...: 3V convergence of ( $m_\beta$ , $m_{\beta\beta}$ , $\Sigma$ ) data

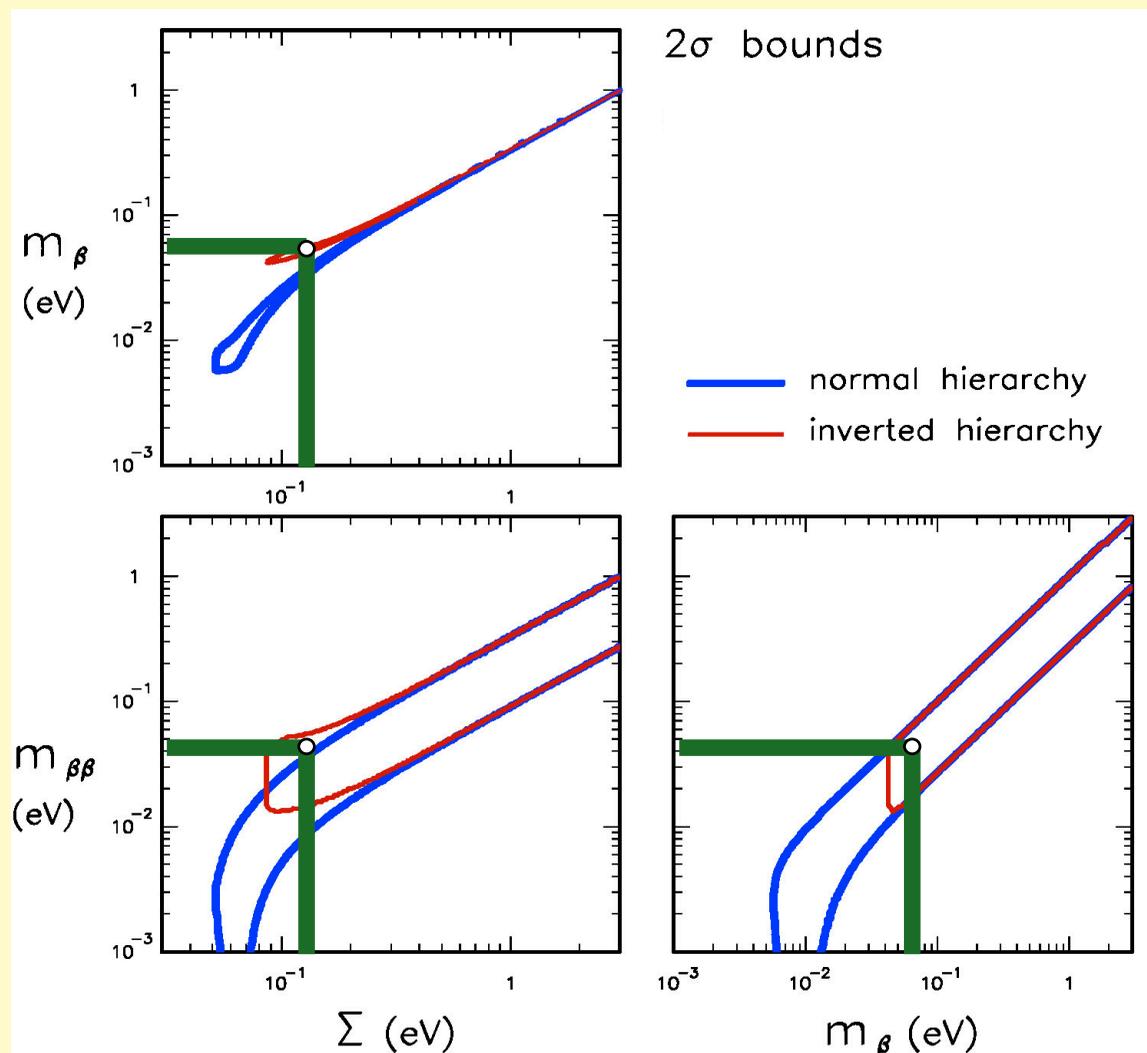
Beta decay  $\nearrow$   $m_\beta$   
 Double beta decay  $\nearrow$   $m_{\beta\beta}$   
 Cosmology  $\nearrow$   $\Sigma$

If realized,  
one might...

Determine the  
mass scale...

Identify the  
hierarchy ...

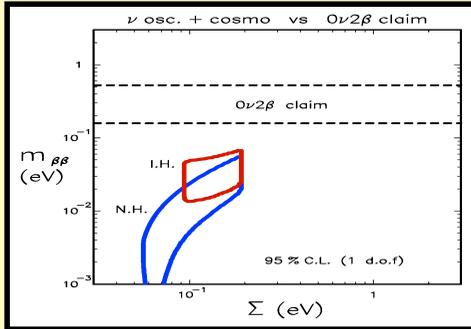
Probe the  
Majorana  
nature and  
phase(s)...



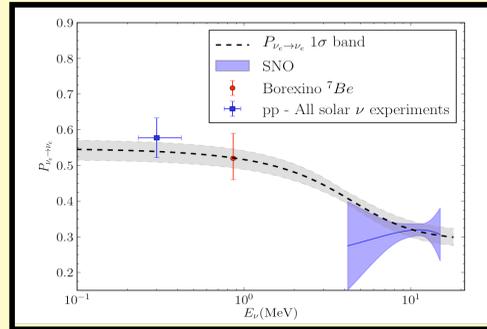
Of course, also “divergences” of results are intriguing, as they might indicate new particles or interactions.

Some more-or-less “anomalous” results still under scrutiny:

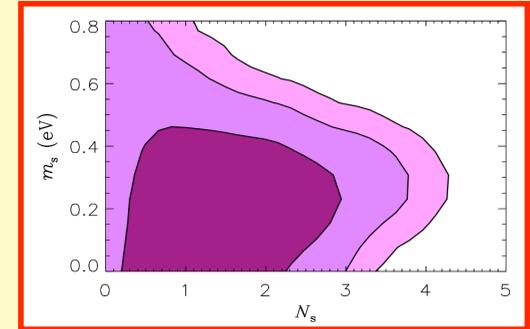
Klapdor claim vs Cosmo?



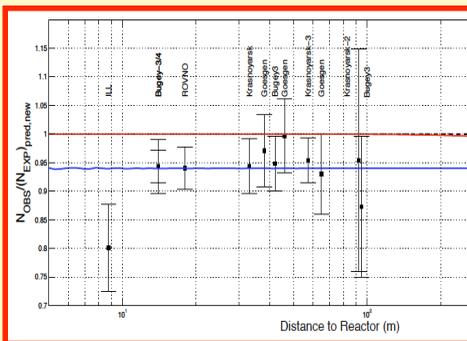
MSW upturn, solar  $^8\text{B}$ ?



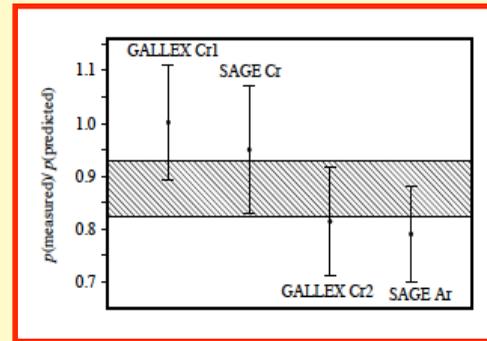
Extra cosmic radiation?



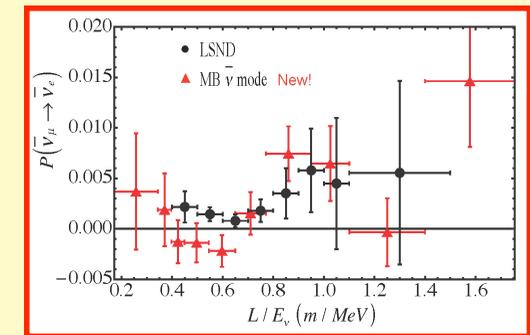
Reactor anomaly?



Gallium anomaly?

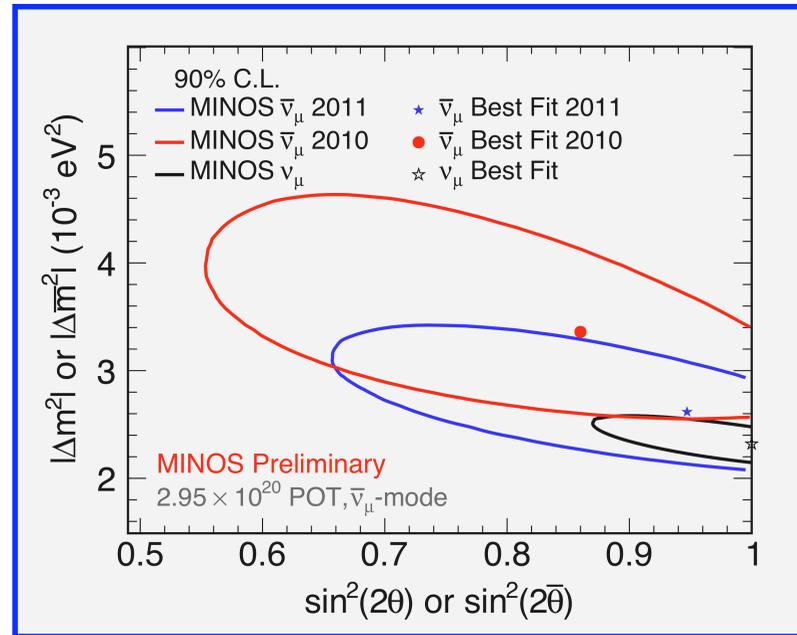


LSND/MiniBooNE (SBL)?



Attempts to interpret the latter four in terms of  $O(\text{eV})$  sterile neutrino(s)

## Footnote: MINOS neutrino/antineutrino anomaly...

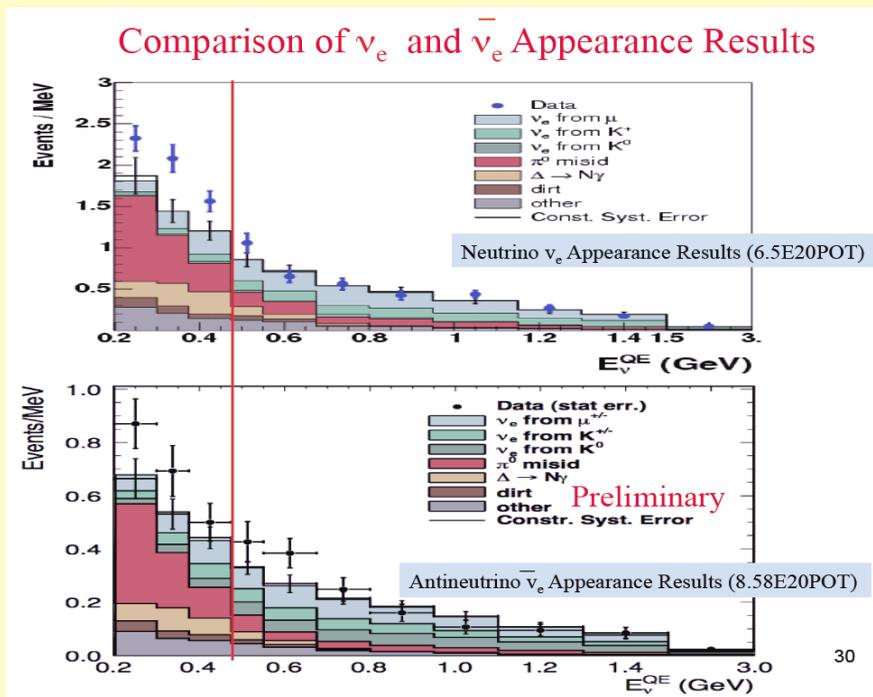


...does not show up in 2011 data!

# QUO VADIS, MiniBooNE?

Reminder: conventional threshold ( $E > 475 \text{ MeV}$ ) initially set to look for LSND-like oscillations in  $\sim$ background-safe range.  
(below: higher background and  $\nu \rightarrow \text{anti-}\nu$  contamination)

2011 data: Similar  $\nu$  & anti- $\nu$  excess, smoothly decreasing with  $E$ .  
Both spectra show no statistically significant signal  $>475 \text{ MeV}$ .



Neutrino excess significance:  
Full E range:  $\sim 3\sigma$   
 $E > 475 \text{ MeV}$ :  $0.6\sigma$

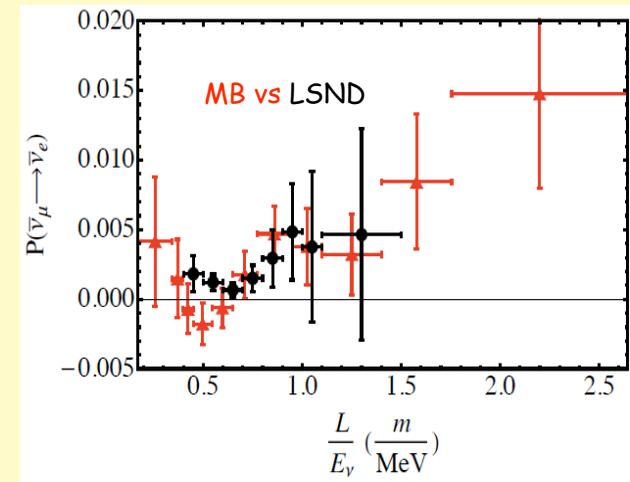
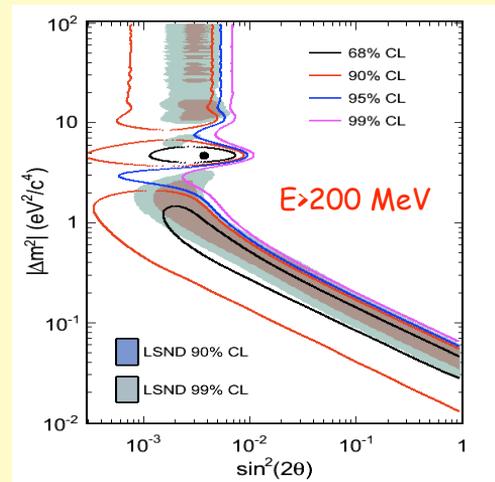
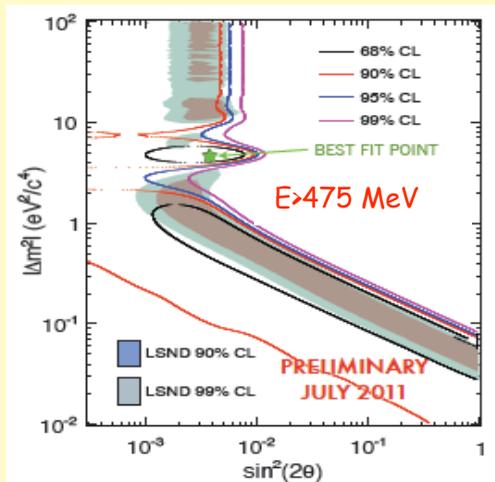
Antineutrino exc. significance:  
Full E range:  $\sim 2\sigma$   
 $E > 475 \text{ MeV}$ :  $0.84\sigma$

## No real indication for LSND-like L/E oscillations...

If the slight MB anti- $\nu$  excess is fitted in terms of  $2\nu$ :

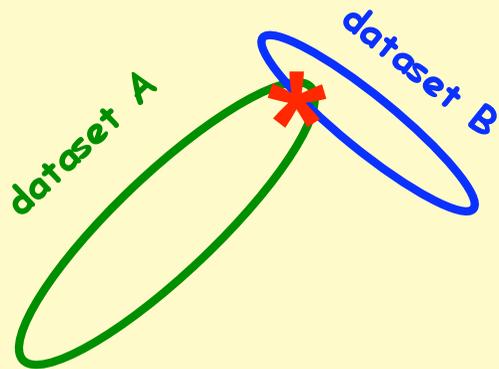
- best fit falls in LSND-excluded region ("out-of-phase");
- at  $\sim 2\sigma$ , both LSND-like and no oscillations are allowed;
- similarly for  $E > 200$  MeV, if  $\nu$  contamination accounted for.

→ "compatibility," rather than "convergence."

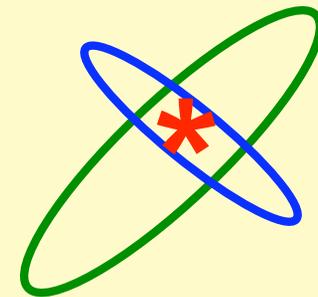


[Preliminary antineutrino plots taken from LSND presentation at PIC 2011]

Similarly, global interpretations of these and other data in terms of **active-sterile** oscillations only show **rough compatibility** among partial data sets, but **not convergence in details**:



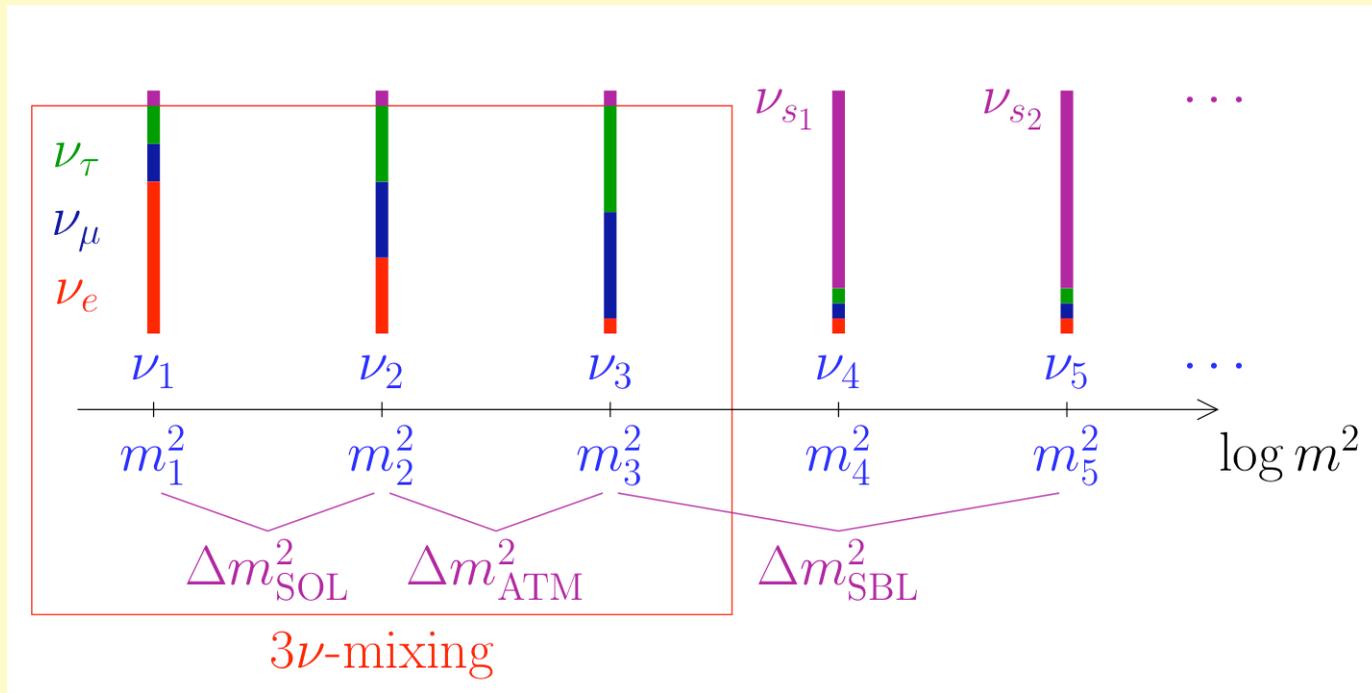
...rather than...



with corresponding **instabilities/poor quality of fits**, living in the "tails" of pdf's or in unstable "corners" of parameter space.

Most disturbing and persistent tension is between LSND (+MB) **appearance** data and world **disappearance** data.

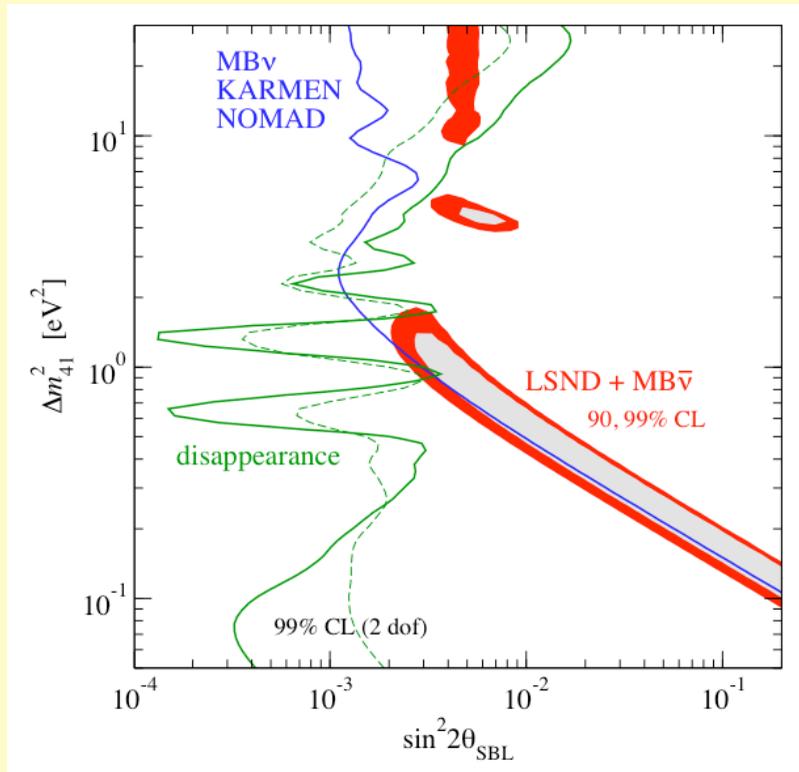
If one or more sterile neutrinos mediate **muon-to-electron appearance** at the LSND L/E scale, then they must induce both **muon and electron disappearance** at the same SBL scale.



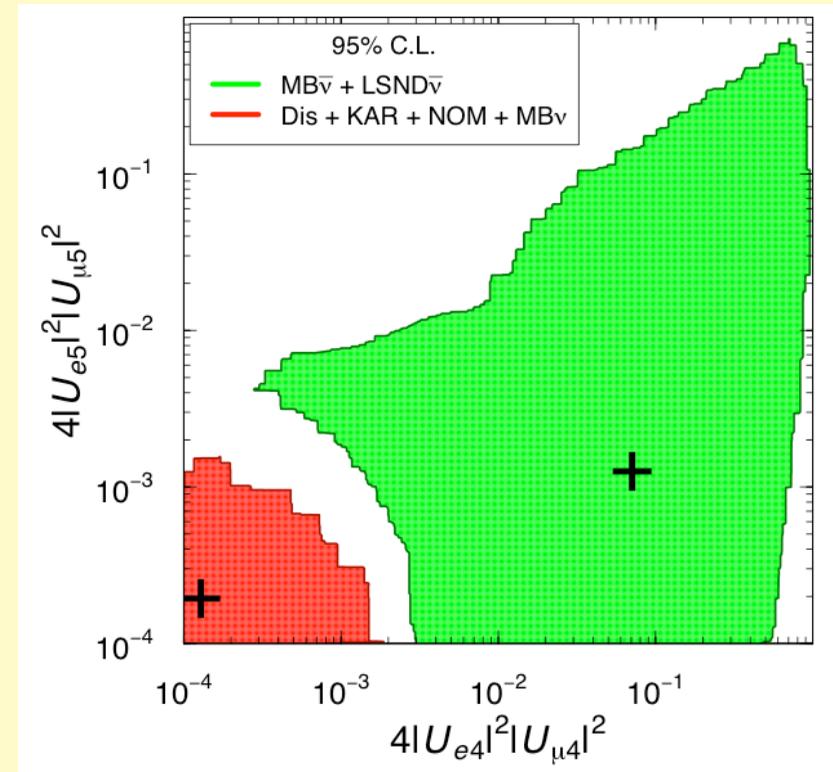
While some electron flavor disappearance is supported by the **reactor and gallium anomalies**, for muon flavor disapp. only upper bounds exist. Constraints do not match  $\rightarrow$

# Appearance/disappearance tension in 3+1 and 3+2

recent fits:

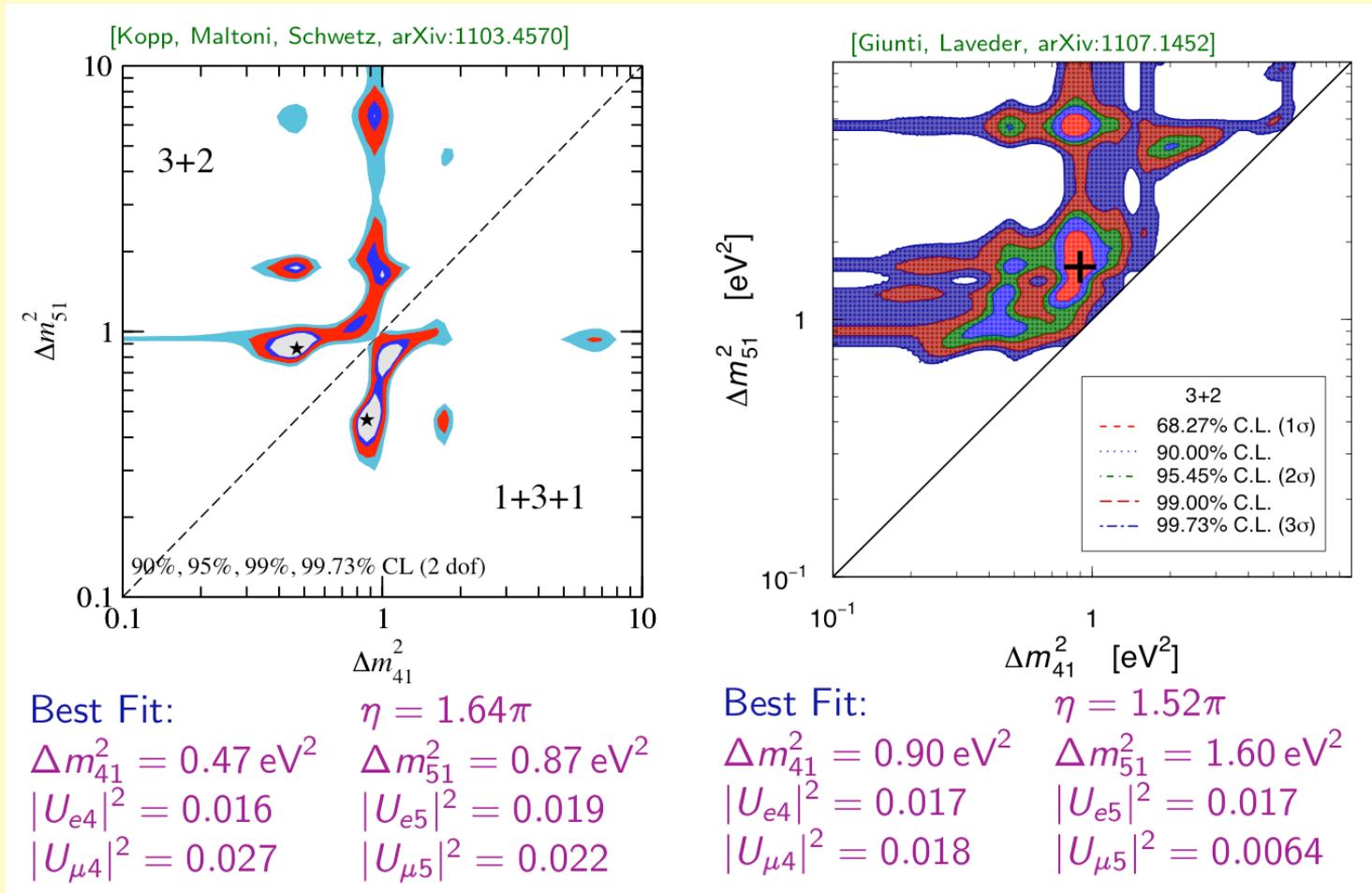


Kopp, Maltoni, Schwetz 2011



Giunti and Laveder 2011

3+2 less strongly disfavored than 3+1, if sterile CP phase tuned...



But then... tension also with standard cosmology: too high  $\nu$  total mass and radiation density [Hamann et al. 2011 + refs. therein]

**Cosmological issues** will be largely clarified before next TAUP  
(relevant Planck release expected early 2013).

**Oscillation issues** are more delicate and call for new accurate expts.  
Using 3+2 best fit as guidance, required sensitivity is challenging:

$$U_{e4}^2 \sim U_{\mu4}^2 \sim U_{e5}^2 \sim U_{\mu5}^2 \sim 0.02$$

→ not easy to test effects of such small mixings

$$\Delta m_{51}^2 \sim O(1) \text{ eV}^2 \sim 2 \times \Delta m_{41}^2$$

→ not easy to disentangle two new frequencies

Moreover, the noted "tensions" suggest that some (or all) of the data may have nothing to do with  $\nu_s$ . More complicated scenarios ( $\nu_s$  + new interactions) can be built, but with "ad hoc" hypotheses.

**Need new, possibly "redundant" expt. tests !**

**Imperative: do not add confusion, aim at  $>3\sigma$  signals !**

## CONCLUSIONS

2011 likely to be another "peak year" in neutrino physics.  
I focused on just two recent topics:

- progress on  $\theta_{13}$  and standard  $3\nu$  parameters,
- renewed interest on SBL data and sterile  $\nu$ 's issues

The standard  $3\nu$  framework remain solid, with converging evidence for  $\theta_{13} > 0$  at  $>3\sigma$ . Precision physics, but still many unknowns: hierarchy, CP, absolute mass, Dirac/Majorana...

There are sparse indications in favor of oscillations mediated by sterile  $\nu$ 's, but a coherent picture has not emerged yet. Even more unknowns in this context...

**Exciting times are guaranteed also for next TAUPs!**